# Self-study manual for introduction to computational fluid dynamics 

## Bachelor's thesis

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Andrey Nabatov

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| Author | Andrey Nabatov | Year 2017 |
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| Supervisor(s) | Esa Murtola |  |


#### Abstract

Computational Fluid Dynamics (CFD) is the branch of Fluid Mechanics and Computational Physics that plays a decent role in modern Mechanical Engineering Design process due to such advantages as relatively low cost of simulation comparing with conduction of real experiment, an opportunity to easily correct the design of a prototype prior to manufacturing of the final product and a wide range of application: mixing, acoustics, cooling and aerodynamics. This makes CFD particularly and Computational Physics in general the tools, desirable to know for every Mechanical Engineering product developer.

This project is aimed to provide the Mechanical Engineering department of HAMK with sufficient study material for a potential course in the future that will introduce HAMK Mechanical Engineering students to Computational Fluid Dynamics. The study material contains parts of theory, describing the ideas behind derivation of governing equations of Fluid Dynamics, their discretization methods, fundamental algorithms of CFD, part of boundary layer theory essential for CFD and turbulence RANS models, which is the required minimum to study prior to the main practical part, consisting of a link to open-source CFD tutorials for ANSYS FLUENT and a simple example done in Mathcad 15, illustrating the work of algorithms and mathematics, covered in theory. There is also an introduction to the Pi-theorem and its applications in the separate chapter, which provides students with a toolbox for a proper analysis of results, gained from several simulations.


Keywords Computational Fluid Dynamics, CFD, study material, instruction, manual.

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## 1 INTRODUCTION

Due to permanently falling cost of computational machinery and its increasing productivity numerical methods for mathematical modelling of complex physics processes are turning to be more economically feasible and commercially attractive techniques in mechanical engineering design work. Therefore, skills of appropriate modelling of physics processes become more and more crucial for employers, when they recruit new workers for mechanical engineering design jobs.

The aim of this thesis is to prepare the study material that can be used for teaching and self-study purposes, when HAMK's Mechanical Engineering department will organize the course/module introducing the students to field of computational physics. Computational Fluid Dynamics were chosen to be the main branch of this work due to several reasons. First, CFD is apparently the only available tool for students and engineers to study complex behaviour of fluids that cannot be described by conventional analytical methods using only pen and paper. Second, CFD is a branch of computational physics that has the same popularity as Finite Element Analysis due to wide variety of areas of usage and their importance: aerodynamics, mixing, cooling, acoustics, combustion, etc.

This work consists of two parts. The first part is theory, covering such topics as governing equations of Fluid Dynamics, discretization methods of governing equations and algorithms of their solving, tips of proper modelling of boundary layer, RANS turbulence models and brief description of principles of Large Eddy Simulation (LES). This theory is a necessary minimum to read prior to start of second part - practicing, containing link to tutorials, done in ANSYS FLUENT, and example, done in Mathcad 15, aimed to show the work of mathematics covered in theoretical part of this thesis and reference to open-source CFD tutorials. Finally, this work contains the key information about Pi-theorem, which is a useful tool for proper analysis of results of multiple experiments and simulations especially in Fluid Mechanics and creating mathematical models based on gathered data with necessary level of accuracy.

## 2 THEORY

### 2.1 Lagrangian versus Eularian approach

In theoretical physics there exist two general approaches describing the motion of fluid flow: Lagrangian and Eularian.
The Lagrangian approach or control mass approach is based on tracking of properties for all mass particles of system as illustrated in Figure 1.


Figure 1. Tracking the path of a single fluid particle, done with Lagrange approach

A good example of an application of the Lagrangian description is the motion of a plane flying from point A to point B. In this case the centre of the mass of the plane is considered to be one single particle with six degrees of freedom traveling on distances several orders higher than size of vehicle. The system has only one mass particle what makes the mathematical model of system to be relatively easy to work with. According to observations of experimental physics, this approach is potentially the closest to reality and mathematically the most accurate in utmost scenarios: when the continuous fluid is equally split to mass particles with size and mass of fluid molecules, appropriate particle collision model and boundary conditions are adjusted. However, this approach is extremely difficult to solve. For instance, in one cubic centimetre of water with 300 K temperature there are approximately $3.33 \times 10^{22}$ water molecules, and every single molecule has to be described by its own Lagrangian function. Finally, it means the need to solve system of $3.33 \times 10^{22} \times 4$ differential equations: one equation per each component of molecule's momentum plus one equation for energy of molecule. Even after applying the numerical approximations and using computational power of modern supercomputers this problem will
remain overcomplicated and computationally expensive. Therefore, the Lagrangian approach usage is currently limited.

The Eularian approach or Control Volume approach on the other hand describes the properties of flow (e.g. velocity, pressure, temperature) as functions of time and space. Instead of modelling the particles of fluid flow, the fluid is assumed to be a continuous medium. In other words, this means that Eularian description requires selecting of the volume field (e.g. space inside the pipe, volume around the plane or car), inside which one wants to determine the properties of flow as shown in Figure 2: the "squeeze" of each spatial cell to infinitely small size will result to analytical solution of fluid transport. And after appropriate setting of boundary conditions and solving governing equations one can obtain flow properties as field functions inside bounded volume. Due to independence of characteristic size from properties of fluid, Eularian description appeared to be the most commonly used method for solving problems of CFD. (www.quora.com 2015).


Figure 2. Visualization of flow as function of time and space done in Eularian approach. The "squeeze" of each spatial cell to infinitely small size will result to analytical solution of fluid transport.

Therefore, theory and computational methods, based on Lagrangian description will be skipped in this thesis project.

### 2.2 Governing equations.

In order to derive the first governing equation, which is continuity equation, one has to consider the infinitesimally small fluid element as shown in Figure 3.


Figure 3. Infinitesimally small volume (control-volume) element in Cartesian coordinate system. (Versteeg \& Malalasekera 2007, $10)$.

The continuity equation states that the change of mass of a single fluid element equals to the difference between flows entering and leaving the flu-id element. The rate of change of mass of fluid element can be expressed as time derivative of density and constant volume of element. Hence change of mass is:

$$
\frac{\partial}{\partial t}(\rho \delta x \delta y \delta z)=\frac{\partial \rho}{\partial t} \delta x \delta y \delta z
$$

To determine the flows entering and leaving the specified control volume, one has to consider three components of mass flow contributing to three different fluid velocity components in three orthogonal directions:

Table 1. Mass flow components due to different fluid velocity components.

| Direction | Velocity <br> component | Mass flow component in exact centre of <br> volume element |
| :---: | :---: | :---: |
| $x$ | $u$ | $\rho u \delta y \delta z$ |
| $y$ | $v$ | $\rho v \delta x \delta z$ |
| $z$ | $w$ | $\rho w \delta x \delta y$ |

Since the volume element has infinitesimally small size, the linear interpolation of flow components with respect to directions of contributing velocity components mathematically will be sufficiently accurate to estimate the flows on volume element faces as illustrated in Figure 4:

Table 2. Mass flow components at centre and faces of volume element.

| Mass flow <br> component at <br> the centre of <br> volume <br> element | Mass flow component <br> entering the volume <br> element | Mass flow component <br> leaving the volume <br> element |
| :---: | :---: | :---: |
| $\rho u \delta y \delta z$ | $\left(\rho u-\frac{\partial(\rho u)}{\partial x} \frac{1}{2} \delta x\right) \delta y \delta z$ | $\left(\rho u+\frac{\partial(\rho u)}{\partial x} \frac{1}{2} \delta x\right) \delta y \delta z$ |
| $\rho v \delta x \delta z$ | $\left(\rho v-\frac{\partial(\rho v)}{\partial y} \frac{1}{2} \delta y\right) \delta x \delta z$ | $\left(\rho v+\frac{\partial(\rho v)}{\partial y} \frac{1}{2} \delta y\right) \delta x \delta z$ |
| $\rho w \delta x \delta y$ | $\left(\rho w-\frac{\partial(\rho w)}{\partial z} \frac{1}{2} \delta z\right) \delta x \delta y$ | $\left(\rho w+\frac{\partial(\rho w)}{\partial z} \frac{1}{2} \delta z\right) \delta x \delta y$ |



Figure 4. Mass flows at faces of volume element (Versteeg \& Malalasekera 2007, 11).

To obtain a flow difference at inlet and outlet faces one has to subtract the summed components at outlet faces from summed components at inlet faces, resulting in:

$$
\begin{aligned}
\left(\rho u-\frac{\partial(\rho u)}{\partial x}\right. & \left.\frac{1}{2} \delta x\right) \delta y \delta z+\left(\rho v-\frac{\partial(\rho v)}{\partial y} \frac{1}{2} \delta y\right) \delta x \delta z \\
& +\left(\rho w-\frac{\partial(\rho w)}{\partial z} \frac{1}{2} \delta z\right) \delta x \delta y-\left(\rho u+\frac{\partial(\rho u)}{\partial x} \frac{1}{2} \delta x\right) \delta y \delta z \\
& -\left(\rho v+\frac{\partial(\rho v)}{\partial y} \frac{1}{2} \delta y\right) \delta x \delta z-\left(\rho w+\frac{\partial(\rho w)}{\partial z} \frac{1}{2} \delta z\right) \delta x \delta y \\
& =-\left(\frac{\partial(\rho u)}{\partial x}+\frac{\partial(\rho v)}{\partial y}+\frac{\partial(\rho w)}{\partial z}\right) \delta x \delta y \delta z
\end{aligned}
$$

Finally, after the setting of equality sign between the rate of change of mass and mass flow difference will result in:

$$
\frac{\partial \rho}{\partial t} \delta x \delta y \delta z=-\left(\frac{\partial(\rho u)}{\partial x}+\frac{\partial(\rho v)}{\partial y}+\frac{\partial(\rho w)}{\partial z}\right) \delta x \delta y \delta z
$$

That after several trivial rearrangements can be simplified to final form of mass continuity equation (Versteeg \& Malalasekera 2007, 9-11):

$$
\begin{equation*}
0=\frac{\partial \rho}{\partial t}+\operatorname{div}(\rho \mathbf{u}) \tag{2.2.1}
\end{equation*}
$$

Where:

$$
\begin{equation*}
\operatorname{div}(\rho \mathbf{u}) \equiv \frac{\partial(\rho u)}{\partial x}+\frac{\partial(\rho v)}{\partial y}+\frac{\partial(\rho w)}{\partial z} \tag{2.2.2}
\end{equation*}
$$

$\rho$ - Is fluid density as a field function of $(x ; y ; z ; t)$.
$u$ - Is component of velocity vector $u$ towards $x$-direction as a field function of $(x ; y ; z ; t)$.
$v$ - Is component of velocity vector $u$ towards $y$-direction as a field function of $(x ; y ; z ; t)$.
$w$ - Is component of velocity vector u towards z -direction as a field function of $(x ; y ; z ; t)$.
$t$ - Is time.

Or:

$$
\mathbf{u}=\left[\begin{array}{l}
u  \tag{2.2.3}\\
v \\
w
\end{array}\right] \begin{aligned}
& (x) \\
& (y) \\
& (z)
\end{aligned}
$$

" $\operatorname{div}(\mathbf{a})$ " is called a divergence of vector " $\mathbf{a}$ " (alternative notation: $\nabla \cdot \mathbf{a}$ ). As an example, for vector a in Cartesian coordinate system:

$$
\begin{gathered}
\mathbf{a}=\left[\begin{array}{l}
a_{x} \\
a_{y} \\
a_{z}
\end{array}\right] \\
\operatorname{div}(\mathbf{a}) \equiv \nabla \cdot \mathbf{a}=\frac{\partial a_{x}}{\partial x}+\frac{\partial a_{y}}{\partial y}+\frac{\partial a_{z}}{\partial z}
\end{gathered}
$$

Equation 2.2.1 is valid for unsteady compressible three-dimensional flow. Further assumption of incompressible, steady flow will lead to next form of continuity differential equation:

$$
\begin{equation*}
0=\frac{\partial u}{\partial x}+\frac{\partial v}{\partial y}+\frac{\partial w}{\partial z} \tag{2.2.4}
\end{equation*}
$$

for three-dimensional incompressible flow, and:

$$
\begin{equation*}
0=\frac{\partial u}{\partial x}+\frac{\partial v}{\partial y} \tag{2.2.5}
\end{equation*}
$$

for two-dimensional incompressible flow.
One can also show that term, equal to zero in eq. 2.2.1 is just one possible form of more generalized term:

$$
\begin{equation*}
\frac{\partial(\rho \varphi)}{\partial t}+\operatorname{div}(\rho \varphi \mathbf{u}) \tag{2.2.6}
\end{equation*}
$$

Where:
$\phi$ - is arbitrary conservative intensive property (e.g. mass, momentum, or energy per unit mass). (Versteeg \& Malalasekera 2007, 12-14).

That combined with eq.2.2.1 can be further simplified to:

$$
\begin{equation*}
\frac{\partial(\rho \varphi)}{\partial t}+\operatorname{div}(\rho \varphi \mathbf{u})=\rho \frac{D \varphi}{D t} \tag{2.2.7}
\end{equation*}
$$

Where:
$\frac{D}{D t}$ - is material derivative operator. Other names of material derivative are advective, convective, hydrodynamic, Lagrangian, particle, substantial, substantive, Stokes or total derivative.

Applying 2.2.7 to the rest conservative properties of flow (momentum and energy), the term $\rho \frac{D \varphi}{D t}$ will turn to $\rho \frac{D u}{D t^{\prime}} \rho \frac{D v}{D t^{\prime}} \rho \frac{D w}{D t}$ for x-momentum, y -momentum and z-momentum respectively, and $\rho \frac{D E}{D t}$ for energy. Note that as rest of $\phi$-properties E is also an intensive property, meaning that in SI units E will have dimensions of $\left[\mathrm{m}^{2} / \mathrm{s}^{2}\right]$ instead of $\left[\mathrm{kg}^{*} \mathrm{~m}^{2} / \mathrm{s}^{2}\right]$. Unlike for continuity, terms $\rho \frac{D u}{D t^{\prime}} \rho \frac{D v}{D t^{\prime}} \rho \frac{D w}{D t}, \rho \frac{D E}{D t}$ are not always equal to 0 . Using the same concept of infinitely small fluid element, it's possible to show that:

$$
\begin{equation*}
\rho \frac{D u}{D t}=\frac{d\left(-p+\tau_{x x}\right)}{d x}+\frac{\partial \tau_{y x}}{\partial y}+\frac{\partial \tau_{z x}}{\partial z}+S_{M x} \tag{2.2.8}
\end{equation*}
$$

$$
\begin{equation*}
\rho \frac{D v}{D t}=\frac{d \tau_{x y}}{d x}+\frac{\partial\left(-p+\tau_{y y}\right)}{\partial y}+\frac{\partial \tau_{z y}}{\partial z}+S_{M y} \tag{2.2.9}
\end{equation*}
$$

$$
\begin{equation*}
\rho \frac{D w}{D t}=\frac{d \tau_{x z}}{d x}+\frac{\partial \tau_{y z}}{\partial y}+\frac{\partial\left(-p+\tau_{y z}\right)}{\partial z}+S_{M z} \tag{2.2.10}
\end{equation*}
$$

Where:
$p$ - is pressure as a field function of $(x ; y ; z ; t)$.
$\tau_{x x}, \tau_{y y}, \tau_{z z}, \tau_{x y}, \tau_{y x}, \tau_{x z}, \tau_{z x}, \tau_{y z}, \tau_{z y}$ - are components of Cauchy viscous stress tensor (see 2.2.11).
$S_{M x}, S_{M y}, S_{M z}$ - are source terms, responsible for effects of other physical phenomena, e.g. if there is a need to consider effects of gravity force, then $S_{M x}=0 ; S_{M y}=0 ; S_{M z}=-\rho g$.
$\tau_{i j}=\left[\begin{array}{lll}\tau_{x x} & \tau_{x y} & \tau_{x z} \\ \tau_{y x} & \tau_{y y} & \tau_{y z} \\ \tau_{z x} & \tau_{z y} & \tau_{z z}\end{array}\right]$ - Cauchy viscous stress tensor, representing
effect of viscosity on fluid motion.

Assuming that fluid is isotropic (all gases in fact are isotropic, only few liquids that contain significant amount of long polymer molecules show anisotropic behaviour) and Newtonian, meaning that $\tau_{i j}$ is linearly proportional to local deformations $s_{i j}$ (see 2.2.12), will give us next relations:

$$
s_{i j}=\left[\begin{array}{ccc}
\frac{\partial u}{\partial x} & \frac{1}{2}\left(\frac{\partial u}{\partial y}+\frac{\partial v}{\partial x}\right) & \frac{1}{2}\left(\frac{\partial u}{\partial z}+\frac{\partial w}{\partial x}\right)  \tag{2.2.12}\\
\frac{1}{2}\left(\frac{\partial u}{\partial y}+\frac{\partial v}{\partial x}\right) & \frac{\partial v}{\partial y} & \frac{1}{2}\left(\frac{\partial v}{\partial z}+\frac{\partial w}{\partial y}\right) \\
\frac{1}{2}\left(\frac{\partial u}{\partial z}+\frac{\partial w}{\partial x}\right) & \frac{1}{2}\left(\frac{\partial v}{\partial z}+\frac{\partial w}{\partial y}\right) & \frac{\partial w}{\partial z}
\end{array}\right]
$$

$$
\begin{equation*}
\tau_{i j}=2 \mu s_{i j} \tag{2.2.13}
\end{equation*}
$$

Where:
$\mu$ - is dynamic viscosity.
Compressibility of fluid can be encountered by introducing the second viscosity " $\lambda$ ", that gives linear relation of stresses to volumetric deformation, equal to $\operatorname{div}(\mathbf{u})$ (see 2.2.14). (Versteeg \& Malalasekera 2007, 14-16).

$$
\begin{equation*}
\operatorname{div}(\mathbf{u}) \equiv \frac{\partial u}{\partial x}+\frac{\partial v}{\partial y}+\frac{\partial w}{\partial z} \tag{2.2.14}
\end{equation*}
$$

Combining of 2.2.12, 2.2.13 and adding effect of compressibility will result in next formulas for Cauchy viscous stress tensor components:

$$
\tau_{x x}=2 \mu \frac{\partial u}{\partial x}+\lambda \operatorname{div}(\mathbf{u})
$$

$$
\tau_{y y}=2 \mu \frac{\partial v}{\partial y}+\lambda \operatorname{div}(\mathbf{u})
$$

$$
\tau_{z z}=2 \mu \frac{\partial w}{\partial z}+\lambda \operatorname{div}(\mathbf{u})
$$

$$
\tau_{x y}=\tau_{y x}=\mu\left(\frac{\partial u}{\partial y}+\frac{\partial v}{\partial x}\right)
$$

$$
\tau_{x z}=\tau_{z x}=\mu\left(\frac{\partial u}{\partial z}+\frac{\partial w}{\partial x}\right)
$$

$$
\begin{equation*}
\tau_{y z}=\tau_{z y}=\mu\left(\frac{\partial v}{\partial z}+\frac{\partial w}{\partial y}\right) \tag{2.2.15}
\end{equation*}
$$

Substituting values of $\tau_{i j}$ from 2.2.15 to 2.2.8, 2.3.9 and 2.2.10 will give:
$\rho \frac{D u}{D t}=-\frac{\partial p}{\partial x}+\frac{\partial}{\partial x}\left[2 \mu \frac{\partial u}{\partial x}+\lambda \operatorname{div}(\mathbf{u})\right]+\frac{\partial}{\partial y}\left[\mu\left(\frac{\partial u}{\partial y}+\frac{\partial v}{\partial x}\right)\right]+\frac{\partial}{\partial z}\left[\mu\left(\frac{\partial u}{\partial z}+\frac{\partial w}{\partial x}\right)\right]+S_{M x}$
(2.2.16)
$\rho \frac{D v}{D t}=-\frac{\partial p}{\partial y}+\frac{\partial}{\partial x}\left[\mu\left(\frac{\partial u}{\partial y}+\frac{\partial v}{\partial x}\right)\right]+\frac{\partial}{\partial y}\left[2 \mu \frac{\partial v}{\partial y}+\lambda \operatorname{div}(\mathbf{u})\right]+\frac{\partial}{\partial z}\left[\mu\left(\frac{\partial v}{\partial z}+\frac{\partial w}{\partial y}\right)\right]+S_{M y}$
(2.2.17)
$\rho \frac{D w}{D t}=-\frac{\partial p}{\partial z}+\frac{\partial}{\partial x}\left[\mu\left(\frac{\partial u}{\partial z}+\frac{\partial w}{\partial x}\right)\right]+\frac{\partial}{\partial y}\left[\mu\left(\frac{\partial v}{\partial z}+\frac{\partial w}{\partial y}\right)\right]+\frac{\partial}{\partial z}\left[2 \mu \frac{\partial w}{\partial z}+\lambda \operatorname{div}(\mathbf{u})\right]+S_{M z}$

There isn't much known about behaviour of " $\lambda$ " as function of various flow and fluid properties. However, there is a suitable approximation for gases exists:

$$
\begin{equation*}
\lambda=-\frac{2}{3} \mu \tag{2.2.19}
\end{equation*}
$$

Combining 2.2.19 with 2.2.16; 2.2.17, 2.2.18 and applying several mathematical rearrangements will result into famous Navier-Stokes equations (Versteeg \& Malalasekera 2007, 21-24):

$$
\begin{equation*}
\rho \frac{D u}{D t}=-\frac{\partial p}{\partial x}+\operatorname{div}(\mu \operatorname{grad}(u))+S_{M x} \tag{2.2.20}
\end{equation*}
$$

$$
\begin{equation*}
\rho \frac{D v}{D t}=-\frac{\partial p}{\partial y}+\operatorname{div}(\mu \operatorname{grad}(v))+S_{M y} \tag{2.2.21}
\end{equation*}
$$

$$
\begin{equation*}
\rho \frac{D w}{D t}=-\frac{\partial p}{\partial z}+\operatorname{div}(\mu \operatorname{grad}(w))+S_{M z} \tag{2.2.22}
\end{equation*}
$$

Where $\operatorname{grad}(\varphi)$ denotes the gradient of scalar property " $\varphi$ ":

$$
\operatorname{grad}(\varphi)=\left[\begin{array}{l}
\frac{\partial \varphi}{\partial x} \\
\frac{\partial \varphi}{\partial y} \\
\frac{\partial \varphi}{\partial z}
\end{array}\right]
$$

If operators of total derivative, divergence and gradient are expanded to form of partial derivatives, Navier-Stokes equations for compressible flow will get next form:

$$
\begin{equation*}
\rho\left(\frac{\partial u}{\partial t}+u \frac{\partial u}{\partial x}+v \frac{\partial u}{\partial y}+w \frac{\partial u}{\partial z}\right)=-\frac{\partial p}{\partial x}+\mu\left(\frac{\partial^{2} u}{\partial x^{2}}+\frac{\partial^{2} u}{\partial y^{2}}+\frac{\partial^{2} u}{\partial z^{2}}\right)-\mu \frac{\partial}{\partial x}\left(\frac{\partial u}{\partial x}+\frac{\partial v}{\partial y}+\frac{\partial w}{\partial z}\right)+S_{M x} \tag{2.2.23}
\end{equation*}
$$

$\rho\left(\frac{\partial v}{\partial t}+u \frac{\partial v}{\partial x}+v \frac{\partial v}{\partial y}+w \frac{\partial v}{\partial z}\right)=-\frac{\partial p}{\partial y}+\mu\left(\frac{\partial^{2} v}{\partial x^{2}}+\frac{\partial^{2} v}{\partial y^{2}}+\frac{\partial^{2} v}{\partial z^{2}}\right)-\mu \frac{\partial}{\partial y}\left(\frac{\partial u}{\partial x}+\frac{\partial v}{\partial y}+\frac{\partial w}{\partial z}\right)+S_{M y}$
(2.2.24)
$\rho\left(\frac{\partial w}{\partial t}+u \frac{\partial w}{\partial x}+v \frac{\partial w}{\partial y}+w \frac{\partial w}{\partial z}\right)=-\frac{\partial p}{\partial z}+\mu\left(\frac{\partial^{2} w}{\partial x^{2}}+\frac{\partial^{2} w}{\partial y^{2}}+\frac{\partial^{2} w}{\partial z^{2}}\right)-\mu \frac{\partial}{\partial z}\left(\frac{\partial u}{\partial x}+\frac{\partial v}{\partial y}+\frac{\partial w}{\partial z}\right)+S_{M z}$
(2.2.25)

For incompressible two-dimensional flow system of equations 2.2.23, 2.2.24, 2.2.25 is simplified to:

$$
\begin{equation*}
\rho\left(\frac{\partial u}{\partial t}+u \frac{\partial u}{\partial x}+v \frac{\partial u}{\partial y}\right)=-\frac{\partial p}{\partial x}+\mu\left(\frac{\partial^{2} u}{\partial x^{2}}+\frac{\partial^{2} u}{\partial y^{2}}\right)+S_{M x} \tag{2.2.26}
\end{equation*}
$$

$$
\begin{equation*}
\rho\left(\frac{\partial v}{\partial t}+u \frac{\partial v}{\partial x}+v \frac{\partial v}{\partial y}\right)=-\frac{\partial p}{\partial y}+\mu\left(\frac{\partial^{2} v}{\partial x^{2}}+\frac{\partial^{2} v}{\partial y^{2}}\right)+S_{M y} \tag{2.2.27}
\end{equation*}
$$

From 2.2.23, 2.2.24, 2.2.25 it's clearly seen that Navier-Stokes equations is a system of second-order partial differential equations that are known to have no universal solution methods. Hence, these equations currently re-main unsolved. However, importance of these equations in science and engineering applications is so huge, that Navier-Stokes equations became the reason to exist for Computational Fluid Dynamics as they are known today and Clay Mathematics Institute offers USD 1000000 reward to one who proves either existence and smoothness, or breakdown of Navier-Stokes equations' solutions.

For last conservative property, that is energy, one can show that governing equation for compressible flow will have a form:

$$
\begin{align*}
& \rho \frac{D E}{D t}=-\operatorname{div}(p \mathbf{u})+\operatorname{div}(k \operatorname{grad}(T)) \\
&+\left[\frac{\partial\left(u \tau_{x x}\right)}{\partial x}+\frac{\partial\left(v \tau_{y x}\right)}{\partial y}+\frac{\partial\left(w \tau_{z x}\right)}{\partial z}+\frac{\partial\left(u \tau_{x y}\right)}{\partial x}+\frac{\partial\left(v \tau_{y y}\right)}{\partial y}+\frac{\partial\left(w \tau_{z y}\right)}{\partial z}\right. \\
&\left.+\frac{\partial\left(u \tau_{x z}\right)}{\partial x}+\frac{\partial\left(v \tau_{y z}\right)}{\partial y}+\frac{\partial\left(w \tau_{z z}\right)}{\partial z}\right]+S_{E} \tag{2.2.28}
\end{align*}
$$

Where:
$E=i+\frac{1}{2}\left(u^{2}+v^{2}+w^{2}\right)-$ sum of kinetic and internal energy.
$i=C_{V} T$
$k$ - is thermal conductivity of fluid.
$T$ - is temperature.
$C_{V^{-}}$is molar heat capacity of gas under constant volume.
Sometimes it might be useful to rearrange 2.2.28, using such properties as internal energy (2.2.29), temperature (2.2.30) or total enthalpy (2.2.31):

$$
\begin{align*}
\rho \frac{D i}{D t}=-p \operatorname{div} & (\mathbf{u})+\operatorname{div}(k \operatorname{grad}(T)) \\
& +\left[\frac{\partial\left(u \tau_{x x}\right)}{\partial x}+\frac{\partial\left(v \tau_{y x}\right)}{\partial y}+\frac{\partial\left(w \tau_{z x}\right)}{\partial z}+\frac{\partial\left(u \tau_{x y}\right)}{\partial x}+\frac{\partial\left(v \tau_{y y}\right)}{\partial y}+\frac{\partial\left(w \tau_{z y}\right)}{\partial z}\right. \\
& \left.+\frac{\partial\left(u \tau_{x z}\right)}{\partial x}+\frac{\partial\left(v \tau_{y z}\right)}{\partial y}+\frac{\partial\left(w \tau_{z z}\right)}{\partial z}\right]+S_{i} \tag{2.2.29}
\end{align*}
$$

$$
\begin{align*}
\rho C_{V} \frac{D T}{D t}=-p & \operatorname{div}(\mathbf{u})+\operatorname{div}(k \operatorname{grad}(T)) \\
& +\left[\frac{\partial\left(u \tau_{x x}\right)}{\partial x}+\frac{\partial\left(v \tau_{y x}\right)}{\partial y}+\frac{\partial\left(w \tau_{z x}\right)}{\partial z}+\frac{\partial\left(u \tau_{x y}\right)}{\partial x}+\frac{\partial\left(v \tau_{y y}\right)}{\partial y}+\frac{\partial\left(w \tau_{z y}\right)}{\partial z}\right. \\
& \left.+\frac{\partial\left(u \tau_{x z}\right)}{\partial x}+\frac{\partial\left(v \tau_{y z}\right)}{\partial y}+\frac{\partial\left(w \tau_{z z}\right)}{\partial z}\right]+S_{i} \tag{2.2.30}
\end{align*}
$$

$$
\begin{align*}
\rho \frac{D h_{0}}{D t}=\frac{\partial p}{\partial t}+ & \operatorname{div}(k \operatorname{grad}(T)) \\
& +\left[\frac{\partial\left(u \tau_{x x}\right)}{\partial x}+\frac{\partial\left(v \tau_{y x}\right)}{\partial y}+\frac{\partial\left(w \tau_{z x}\right)}{\partial z}+\frac{\partial\left(u \tau_{x y}\right)}{\partial x}+\frac{\partial\left(v \tau_{y y}\right)}{\partial y}+\frac{\partial\left(w \tau_{z y}\right)}{\partial z}\right. \\
& \left.+\frac{\partial\left(u \tau_{x z}\right)}{\partial x}+\frac{\partial\left(v \tau_{y z}\right)}{\partial y}+\frac{\partial\left(w \tau_{z z}\right)}{\partial z}\right]+S_{h} \tag{2.2.31}
\end{align*}
$$

Where:

$$
h=i+\frac{p}{\rho} \text { - is enthalpy of gas. }
$$

$h_{0}=h+\frac{1}{2}\left(u^{2}+v^{2}+w^{2}\right)-$ is total enthalpy.

Application of Newtonian viscosity model to eq.2.2.29 will yield to:

$$
\begin{equation*}
\rho C_{V} \frac{D T}{D t}=-p \operatorname{div}(\mathbf{u})+\operatorname{div}(k \operatorname{grad}(T))+\Phi+S_{i} \tag{2.2.32}
\end{equation*}
$$

Where:
$\Phi=\mu\left\{2\left[\left(\frac{\partial u}{\partial x}\right)^{2}+\left(\frac{\partial v}{\partial y}\right)^{2}+\left(\frac{\partial w}{\partial z}\right)^{2}\right]+\left(\frac{\partial u}{\partial y}+\frac{\partial v}{\partial x}\right)^{2}+\left(\frac{\partial u}{\partial z}+\frac{\partial w}{\partial x}\right)^{2}+\left(\frac{\partial v}{\partial z}+\frac{\partial w}{\partial y}\right)^{2}\right\}+\lambda(\operatorname{div}(\mathbf{u}))^{2}$

Together equations 2.2.1, 2.2.20, 2.2.21, 2.2.22, 2.2.29 and such equations of state as ideal gas equation ( $p V=n R T$ ) and internal energy equation ( $i=C_{V} T$ ) are forming the system of seven equations with seven unknowns, meaning that system is mathematically closed (it can be solved, providing that initial and boundary conditions are stated). (Versteeg \& Malalasekera 2007, 18-21).
2.2.1, 2.2.20, 2.2.21, 2.2.22, 2.2.29 can be generalized, using arbitrary property $\varphi$ :

$$
\begin{equation*}
\frac{\partial(\rho \varphi)}{\partial t}+\operatorname{div}(\rho \varphi \mathbf{u})=\operatorname{div}(\Gamma \operatorname{grad}(\varphi))+S_{\varphi} \tag{2.2.34}
\end{equation*}
$$

Where $\Gamma$ - is diffusion coefficient such as viscosity $\mu$, or conductivity $k$.

The eq.2.2.34 is called transport equation of property $\varphi$, which can be changed back to equations 2.2.1, 2.2.20, 2.2.21, 2.2.22, 2.2 .29 by setting the property $\varphi$ to be equal to $1, u, v, w$ or $i$ respectively and appropriately setting of source terms and values of $\Gamma$. In order to get form of eq.2.2.34, more convenient for finite volume method, the integration over Control volume must be applied to eq.2.2.34:

$$
\begin{equation*}
\int_{C V} \frac{\partial(\rho \varphi)}{\partial t} d V+\int_{C V} \operatorname{div}(\rho \varphi \mathbf{u}) d V=\int_{C V} \operatorname{div}(\Gamma \operatorname{grad}(\varphi)) d V+\int_{C V} S_{\varphi} d V \tag{2.2.35}
\end{equation*}
$$

Where:
$V$ - is volume.
CV - is reference to control volume.
CS - is reference to control surface (boundary of control volume).

For further transforming of 2.2.35, there is a need to introduce Gauss's divergence theorem, stating, that for vector " $\mathbf{a}$ ":

$$
\begin{equation*}
\int_{C V} \operatorname{div}(\mathbf{a}) d V=\int_{\mathrm{CS}} \mathbf{n} \cdot \mathbf{a} d A \tag{2.2.36}
\end{equation*}
$$

Where:
$A$ - is area.
$\mathbf{n} . \mathbf{a}$ - is component of vector " $\mathbf{a}$ " in the direction of unit vector " $\mathbf{n}$ ", normal to surface element $d A$.

Applying Gauss's divergence theorem (2.2.36) and changing the order of integration and differentiation in first term in left-hand side of eq.2.2.35 will yield to:

$$
\begin{equation*}
\frac{\partial}{\partial t}\left(\int_{C V}(\rho \varphi) d V\right)+\int_{C S} \mathbf{n} \cdot(\rho \varphi \mathbf{u}) d A=\int_{C S} \mathbf{n} \cdot(\Gamma \operatorname{grad}(\varphi)) d A+\int_{C V} S_{\varphi} d V \tag{2.2.37}
\end{equation*}
$$

To derive the most general integrated form of transport equation, the integration over time interval from t to $t+\Delta t$ must be applied in order to cover time-dependant problems. (Versteeg \& Malalasekera 2007, 24 26):

$$
\int_{\Delta t} \frac{\partial}{\partial t}\left(\int_{C V}(\rho \varphi) d V\right) d t+\int_{\Delta t} \int_{C S} \mathbf{n} \cdot(\rho \varphi \mathbf{u}) d A d t=\int_{\Delta t} \int_{C S} \mathbf{n} \cdot(\Gamma \operatorname{grad}(\varphi)) d A d t+\int_{\Delta t} \int_{C V} S_{\varphi} d V d t
$$

(2.2.38)

### 2.3 Turbulence

In spite the fact that all conservative properties were encountered in subchapter 2.2 the current mathematical model is sufficient only for laminar flows. Various experiments such as ink flow in pipe, illustrated in Figure 5, or a flow past a sphere, shown in Figure 6 and 7, clearly show that at some critical value of fluid velocity in inlet of flow domain the fluid flow starts to generate rotational structures, called eddies, of various scale, causing it to behave in a very unstable and chaotic manner. These chaotic fluctuations of fluid flow parameters, responsible for high energy losses due to conversion of kinetic energy of eddies to internal energy and its further dissipation are called turbulence: from Latin word turbulentus chaotic. A corresponding flow is called a turbulent flow. Particularly in Figure 5 one can see that increasing of inlet velocity yields to increased ink velocity fluctuations along the path of flow (from left to right) and, finally, to mixing with water. Therefore, high turbulence is a compulsory constraint for problems of mixing in engineering: higher turbulence means better mixing. Oppositely, the flow, where such eddies are absent, is called a laminar flow.


Figure 5. Experiment with ink in a water flow.


Figure 6. Laminar flow past sphere.


Figure 7. Turbulent flow past sphere.
The turbulence of a flow is characterised by a non-dimensional parameter, called "Reynolds number":

$$
\begin{equation*}
R e=\frac{\rho U_{r e f} L_{c}}{\mu} \tag{2.3.1}
\end{equation*}
$$

Where:
$R e$ - is Reynolds number.
$\rho$ - is fluid density.
$L_{c}$ - is characteristic length of problem such as diameter or length of pipe or any other linear geometric parameter, the control volume is characterised by.
$\mu$ - is dynamic viscosity of fluid in flow inlet.
$U_{r e f}$ - is the reference mean flow velocity.

The exceed of some critical value of Reynolds number will result to transformation of flow from laminar to turbulent. This critical value isn't common for all problems and has a wide variation range. For instance for flow in pipe the critical value of Reynolds number is about 2300, while for problem of flow past circular object (e.g. cylinder, sphere) the critical Reynolds number value will be about 47.5.

One of the most practically useful but not the only method to describe turbulent flows is usage of Reynolds decomposition, that is a basis for all RANS (Reynolds-Averaged-Navier-Stokes) turbulence models.
Reynolds decomposition method states, that any property of turbulent fluid flow, say velocity $u(t)$, can be decomposed to steady mean value $U$ and fluctuating $u^{\prime}(t)$ components as illustrated in Figure 8, or:

$$
\begin{equation*}
u(t)=U+u^{\prime}(t) \tag{2.3.2}
\end{equation*}
$$



Figure 8. Time-averaged velocity " $u$ " as a part of RANS modelling approach (Versteeg \& Malalasekera 2007, 41).

This approach allows to describe the turbulent flow in terms of timeaveraged properties (like U) and statistical values of their fluctuating parts such as $u^{\prime}(t)$. More generally:

$$
\begin{equation*}
\varphi(t)=\Phi+\varphi^{\prime}(t) \tag{2.3.3}
\end{equation*}
$$

Where:

$$
\begin{equation*}
\Phi=\frac{1}{\Delta t} \int_{0}^{\Delta t} \varphi(t) d t \tag{2.3.4}
\end{equation*}
$$

And:

$$
\begin{equation*}
\frac{1}{\Delta t} \int_{0}^{\Delta t} \varphi^{\prime}(t) d t=0 \tag{2.3.5}
\end{equation*}
$$

To describe the spread of fluctuations $\varphi^{\prime}(t)$ about the mean value $\Phi$, one can use statistical parameters like a variance (2.3.6), root mean square (r.m.s.) (2.3.7), or even higher order moments like a skewness (asymmetry) (2.3.8) and a kurtosis (peakedness) (2.3.9). (Versteeg \& Malalasekera 2007, 49-52).

$$
\begin{equation*}
\overline{\left(\varphi^{\prime}\right)^{2}}=\frac{1}{\Delta t} \int_{0}^{\Delta t}\left(\varphi^{\prime}\right)^{2} d t \tag{2.3.6}
\end{equation*}
$$

$$
\begin{equation*}
\varphi_{r m s}=\sqrt{\overline{\left(\varphi^{\prime}\right)^{2}}}=\left[\frac{1}{\Delta t} \int_{0}^{\Delta t}\left(\varphi^{\prime}\right)^{2} d t\right]^{\frac{1}{2}} \tag{2.3.7}
\end{equation*}
$$

$$
\begin{equation*}
\overline{\left(\varphi^{\prime}\right)^{3}}=\frac{1}{\Delta t} \int_{0}^{\Delta t}\left(\varphi^{\prime}\right)^{3} d t \tag{2.3.8}
\end{equation*}
$$

$$
\begin{equation*}
\overline{\left(\varphi^{\prime}\right)^{4}}=\frac{1}{\Delta t} \int_{0}^{\Delta t}\left(\varphi^{\prime}\right)^{4} d t \tag{2.3.9}
\end{equation*}
$$

One can show that there exists infinite amount of statistical parameters by taking the $\varphi^{\prime}$ to any power " $n$ " and further averaging of $\left(\varphi^{\prime}\right)^{n}$,
however, except variance and root mean square, all these parameters are rarely used in today turbulence modelling problems. On the other hand, variance has a straight connection to total kinetic energy of turbulence " $k$ " per unit mass at certain point:

$$
\begin{equation*}
k=\frac{1}{2}\left[\overline{\left(u^{\prime}\right)^{2}}+\overline{\left(v^{\prime}\right)^{2}}+\overline{\left(w^{\prime}\right)^{2}}\right] \tag{2.3.10}
\end{equation*}
$$

The turbulence intensity $T_{i}$, which is another important parameter for RANS turbulence models often used for boundary condition specification in CFD codes, is also related with velocity variance by linkage with " $k$ ":

$$
\begin{equation*}
T_{i}=\frac{\left(\frac{2}{3} k\right)^{\frac{1}{2}}}{U_{r e f}} \tag{2.3.1}
\end{equation*}
$$

Where term $\left(\frac{2}{3} k\right)^{\frac{1}{2}}$ is indeed an average r.m.s. velocity of fluid.
The variance is also called the second moment of the fluctuations (similarly third and fourth moments of fluctuations for skewness and kurtosis respectively). Important information about a fluid flow is also contained in moments, constructed from two different variables. As an example one can consider two arbitrary properties $\varphi$ and $\psi$. Applying 2.3.3, one defines the second moment of $\varphi^{\prime}$ and $\psi^{\prime}$ as follows:

$$
\begin{equation*}
\overline{\varphi^{\prime} \psi^{\prime}}=\frac{1}{\Delta t} \int_{0}^{\Delta t} \varphi^{\prime} \psi^{\prime} d t \tag{2.3.12}
\end{equation*}
$$

Such second moments are especially crucial in RANS turbulence due to their usage in description of an additional shear stress experienced by fluid in turbulent flow. Another application of second moments are autocorrelation functions, used to study relations between fluctuations at different time instants and space points. Autocorrelation functions are defined as follows:

$$
R_{\varphi^{\prime} \varphi^{\prime}}(\tau)=\overline{\varphi^{\prime}(t) \varphi(t+\tau)}=\frac{1}{\Delta t} \int_{0}^{\Delta t} \varphi^{\prime}(t) \varphi^{\prime}(t+\tau) d t
$$

(2.3.13) (autocorrelation function for different time instants)

$$
R_{\varphi^{\prime} \varphi^{\prime}}(\xi)=\overline{\varphi^{\prime}(\boldsymbol{x} ; t) \varphi^{\prime}(\mathbf{x}+\xi ; t)}=\frac{1}{\Delta t} \int_{t}^{t+\Delta t} \varphi^{\prime}\left(\mathbf{x} ; t^{\prime}\right) \varphi^{\prime}\left(\mathbf{x}+\xi ; t^{\prime}\right) d t^{\prime}
$$

(2.3.14) (autocorrelation function for two points displaced by vector $\pm \xi$ from each other)

Where:
$\tau$-is a time shift constant.
$\mathbf{x}=\mathbf{x}(x ; y ; z)$ - is a shorter notation for position vector, dependent on $x$, $y$ and $z$.

It can be easily checked if either $\tau$ in 2.3 .13 or $|\xi|$ in 2.3 .14 are equal to zero, the correlation function will turn variance, that is told to be perfectly correlated, and will have the largest possible value as function of $\tau$ or $\xi$. Therefore, as $\tau$ or $|\xi|$ approach infinity, the correlation function will decrease to zero. This makes autocorrelation functions to be a useful tool for description of eddy size and lifetime. The integral time and scale, which represent concrete values of average period or size of a turbulent eddy, can be computed from integrals of functions $R_{\varphi^{\prime} \varphi^{\prime}}(\tau)$ with respect to $\tau$ and $R_{\varphi^{\prime} \varphi^{\prime}}(\xi)$ with respect to distance in the direction of one of components of displacement vector $\xi$. By analogy, it is also possible to define cross-correlation functions $R_{\varphi^{\prime} \varphi^{\prime}}(\tau)$ with respect to $\tau$ or $R_{\varphi^{\prime} \varphi^{\prime}}(\xi)$ between pairs of different fluctuations by replacing second $\varphi^{\prime}$ by $\psi^{\prime}$ in equations 2.3.13 and 2.3.14 respectively. (Versteeg \& Malalasekera 2007, 49-52).

### 2.3.1 The law of the wall

As one previously stated, general solution for governing equations of fluid mechanics remain unfound, limiting engineers and scientists with analytic solutions of several simple laminar flow problems. Therefore, due to higher mathematical complexity there exist even less models suitable for turbulent flows. One of such models is called "Law of the wall" which is practically useful for accurate estimation of first mesh cell height from the solid wall. This law plays important role in CFD modelling, originating from no-slip conditions (fluid velocity at wall surface equals zero), which result in high velocity gradients at near-wall region and formation of boundary layer. This means that the grid (mesh) at near-wall regions must be much finer comparing the rest of flow domain, in order to simulate boundary layer profile and other coupled properties (e.g. pressure, temperature etc.) And the law of the wall so far remains to be the best tool to encounter these crucial aspects.

In order to formulate the law of the wall one needs to introduce two other non-dimensional parameters:
$u^{+}=\frac{u}{u_{\tau}}-$ called dimensionless velocity, and
$y^{+}=\frac{\rho y u_{\tau}}{\mu}-$ is dimensionless wall coordinate

Where:
$u$ - is a fluid velocity, parallel to the wall,
$u_{\tau}=\sqrt{\frac{\tau_{w}}{\rho}}$ is a friction velocity,
$\tau_{w}$ - is a viscous shear stress,
$y$ - is a distance coordinate, normal to wall surface,
$\rho$ - is a fluid density,
$\mu$ - is dynamic viscosity.
The law of the wall itself states the relation between two parameters $u^{+}$ and $y^{+}$, forming $u^{+}$as function of $y^{+}\left(u^{+}=\mathrm{f}\left(y^{+}\right)\right)$for high Reynolds numbers in a next form:

1. For any $y^{+}<5$ the fluid flow is in region of viscous sublayer of flow boundary layer, characterised by laminar behaviour of flow due to 0 fluid velocity at the level of wall, which is a consequence of fluid property to stick to the wall of solid and almost constant value of $\tau_{w}$. In viscous sublayer the next relation between $u^{+}$and $y^{+}$ holds:

$$
\begin{equation*}
u^{+}=y^{+} \tag{2.3.18}
\end{equation*}
$$

2. For $5<y^{+}<30$ the flow is part of buffer layer which can be described with certain error by both laws from previous section and from next one.
3. For $30<y^{+}<500$ the next expression is valid:

$$
\begin{equation*}
u^{+}=\frac{1}{\kappa} \ln \left(y^{+}\right)+C^{+} \tag{2.3.19}
\end{equation*}
$$

Where: $\kappa=0.4187, C^{+}=5.1$ are constants. Note that they are valid only for smooth walls, the most common case in CFD. For more details see Schlichting, H. (1979) Boundary-layer Theory.

The part that has the most useful information is contained in first section of stated law. The expression $u^{+}=y^{+}$for any $y^{+}<5$ implies that the velocity profile inside viscous sublayer shows the linear behaviour with respect to distance from the wall. This means that in CFD modelling problems, where linear approximations are key for solving fluid dynamics problems, it is sufficient to use just 1 volume element for complete description of flow inside the viscous sublayer. Therefore, the law of the wall serves as the answer for the problem of first mesh cell height calculation, allowing CFD programme users to find the height of near-wall volume elements that will be enough for accurate modelling of the boundary layer on solid walls.

Finally the first mesh cell height can be expressed as function of Reynolds number (or it's individual parameters: free-stream velocity, fluid density, dynamic viscosity and reference length) and $y^{+}$.

Links to first cell height on-line calculators:
https://www.computationalfluiddynamics.com.au/tips-tricks-cfd-
estimate-first-cell-height/
https://geolab.larc.nasa.gov/APPS/YPlus/
http://www.pointwise.com/yplus/
https://www.cfd-online.com/Tools/yplus.php
Note, that all calculations are based on one particular fluid dynamics problem of flat-plate boundary layer (or pipe-channel flow), where there is an only one option for value of reference length that is the length of plate. The vast majorities of fluid dynamics problems, showing poor similarity to flat-plate boundary layer problem, involve geometries that are dependent on multiple linear parameters (e.g. length, width, height, rounding radius etc.), and all of them can be treated as reference lengths. This means that usage of different reference lengths in grid-spacing calculators will yield to different values of mesh cell height, which is not acceptable. Therefore, in order to ensure that cell height is sufficiently small but still relevant to particular problem, one should use desired " $y^{+"}$ value to be less or equal to 1 , enhanced wall treatment must be enabled and there is sufficient amount of cells to resolve the whole boundary layer.

Because mentioned instruction does not always guarantee success in proper modelling of viscous sublayer one might need to find suitable near-wall cell height empirically, by gradual refinement of near-wall mesh after each simulation. This approach, however, is suggested to be used only as last step after previous methods failed in viscous sublayer modelling.

Alternatively, in order to save computational time one can use so-called wall-functions. Similarly to first cell height calculation, background of wall
functions is the same law of the wall. However, instead of computing the wall adjacent height for viscous sublayer, one has to compute the first cell height for the whole boundary layer. Particularly scalable wall function requires the corresponding value of $y^{+}=11.225$, and in case of wrong estimation, the programme will shift the height of first cell to this value automatically. Therefore, the usage of wall functions allows to use much coarser grids, what makes them extremely popular in industrial applications. On the other hand, comparing with first approach, wall functions have two drawbacks:

- Simulation results with enabled wall functions and coarse mesh are less accurate than results with fine mesh and disabled wall functions.
- Wall functions are not applicable for cases with flow separation as shown in Figure 9.


Figure 9. Wall functions are not applicable to problems involving a flow separation.

More information regarding the " $y^{+"}$ and first cell height estimation can be found via following link:
https://www.computationalfluiddynamics.com.au/tag/wall-functions/
2.3.2 Introduction to RANS models

Recalling formulas $2.3 .3,2.3 .4,2.3 .5,2.3 .6,2.3 .7$ and 2.3 .12 one can show that next expressions for derivatives and integrals for arbitrary scalar properties $\varphi$ and $\psi$ hold:

$$
\begin{gathered}
\overline{\varphi^{\prime}}=\overline{\psi^{\prime}}=0 \\
\bar{\Phi}=\Phi \\
\frac{\overline{\partial \varphi}}{\partial s}=\frac{\partial \Phi}{\partial s}
\end{gathered}
$$

$$
\begin{gather*}
\overline{\int \varphi \varphi d s}=\int \Phi \mathrm{ds} \\
\hline \overline{\varphi+\psi}=\Phi+\Psi \\
\overline{\varphi \bar{\psi}}=\Psi \Phi+\overline{\varphi^{\prime} \psi^{\prime}} \\
\hline \overline{\varphi \Psi}=\Phi \Psi \\
\overline{\varphi^{\prime} \Psi}=0 \tag{2.3.20}
\end{gather*}
$$

Since gradient and divergence are both differentiation operators, next expressions for arbitrary vector quantity $\mathbf{a}=\mathbf{A}+\mathbf{a}^{\prime}$ and arbitrary scalar quantity $\varphi=\Phi+\varphi^{\prime}$ can also be proven to be valid:

$$
\begin{gather*}
\overline{\operatorname{div}(\mathbf{a})}=\operatorname{div}(\mathbf{A}) \\
\overline{\operatorname{div}(\varphi \mathbf{a})}=\operatorname{div}(\varphi \mathbf{a})=\operatorname{div}(\Phi \mathbf{A})+\operatorname{div}\left(\overline{\varphi^{\prime} \mathbf{a}^{\prime}}\right) \\
\overline{\operatorname{div}(\operatorname{grad}(\varphi))}=\operatorname{div}(\operatorname{grad}(\Phi)) \tag{2.3.21}
\end{gather*}
$$

Substitution of formulas 2.3.3, 2.3.20, 2.3.21 to governing equations 2.2.1, 2.2.20, 2.2.21, 2.2.22 and 2.2.34, combined with Favre averaging (2.3.22) will yield to averaged turbulent flow equations also called as RANS equations model (where RANS stands for Reynolds-averaged Navier-Stokes) (2.3.23). (Versteeg \& Malalasekera 2007, 62).

$$
\begin{gather*}
\varphi=\widetilde{\Phi}+\varphi^{\prime \prime} \\
\widetilde{\Phi}=\frac{\int_{\Delta t} \rho(t) \varphi(t)}{\int_{\Delta t} \rho(t)}=\frac{\overline{\rho \varphi}}{\bar{\rho}} \\
\overline{\rho \varphi^{\prime \prime}}=0 \\
\overline{\rho \widetilde{\Phi}}=\bar{\rho} \widetilde{\Phi}=\overline{\rho \varphi} \tag{2.3.22}
\end{gather*}
$$

RANS equations:
Continuity:

$$
\frac{\partial \bar{\rho}}{\partial t}+\operatorname{div}(\bar{\rho} \widetilde{\mathbf{U}})=0
$$

Reynolds equations:

$$
\begin{aligned}
& \frac{\partial(\bar{\rho} \widetilde{U})}{\partial t}+\operatorname{div}(\bar{\rho} \widetilde{U} \widetilde{\mathbf{U}})=-\frac{\partial P}{\partial x}+\operatorname{div}(\mu \operatorname{grad}(\widetilde{U}))+\left[-\frac{\partial \overline{\left(\bar{\rho} u^{\prime 2}\right)}}{\partial x}-\frac{\partial \overline{\left(\bar{\rho} u^{\prime} v^{\prime}\right)}}{\partial y}-\frac{\partial \overline{\left(\bar{\rho} u^{\prime} w^{\prime}\right)}}{\partial z}\right]+S_{M x} \\
& \left.\frac{\partial(\bar{\rho} \tilde{V})}{\partial t}+\operatorname{div}(\bar{\rho} \tilde{V} \widetilde{\mathbf{U}})=-\frac{\partial P}{\partial y}+\operatorname{div}(\mu \operatorname{grad}(\tilde{V}))+\left[-\frac{\partial \overline{\left(\bar{\rho} u^{\prime} v^{\prime}\right)}}{\partial x}-\frac{\partial \overline{\left(\bar{\rho} v^{\prime 2}\right)}}{\partial y}-\frac{\partial \overline{\left(\bar{\rho} v^{\prime} w^{\prime}\right)}}{\partial z}\right]+S_{M y}\right] \\
& \frac{\partial(\bar{\rho} \widetilde{W})}{\partial t}+\operatorname{div}(\bar{\rho} \widetilde{W} \widetilde{\mathbf{U}})=-\frac{\partial P}{\partial z}+\operatorname{div}(\mu \operatorname{grad}(\widetilde{W}))+\left[-\frac{\partial \overline{\left(\bar{\rho} u^{\prime} w^{\prime}\right)}}{\partial y}-\frac{\partial \overline{\left(\bar{\rho} v^{\prime} w^{\prime}\right)}}{\partial y}-\frac{\partial \overline{\left(\bar{\rho} w^{\prime 2}\right)}}{\partial z}\right]+S_{M z}
\end{aligned}
$$

Scalar transport equation:

$$
\begin{equation*}
\frac{\partial(\bar{\rho} \widetilde{\Phi})}{\partial t}+\operatorname{div}(\bar{\rho} \widetilde{\Phi} \widetilde{\mathbf{U}})=\operatorname{div}\left(\Gamma_{\Phi} \operatorname{grad}(\widetilde{\Phi})\right)+\left[-\frac{\partial \overline{\left(\bar{\rho} u^{\prime} \varphi^{\prime}\right)}}{\partial x}-\frac{\partial \overline{\left(\bar{\rho} v^{\prime} \varphi^{\prime}\right)}}{\partial y}-\frac{\partial \overline{\left(\bar{\rho} w^{\prime} \varphi^{\prime}\right)}}{\partial z}\right]+S_{\Phi} \tag{2.3.23}
\end{equation*}
$$

Where:
$\boldsymbol{u}=\mathbf{U}+\boldsymbol{u}^{\prime}=\widetilde{\mathbf{U}}+\boldsymbol{u}^{\prime \prime} ;$
$u=U+u^{\prime}=\widetilde{U}+u^{\prime \prime} ;$
$v=V+v^{\prime}=\tilde{V}+v^{\prime \prime} ;$
$w=W+w^{\prime}=\widetilde{W}+w^{\prime \prime} ;$
$p=P+p^{\prime}$;
The terms in square brackets in Reynolds equations in 2.3.23 are extra turbulence stresses, also called Reynolds stresses. They are always nonzero and very large, comparing with the viscous stresses in a turbulent flow. Therefore, neglecting of these terms can cause poor accuracy of final solution, making the simulation results completely irrelevant to experimental data. (Versteeg \& Malalasekera 2007, 65).

Equations 2.3.23 introduce 6 extra unknowns to the system of governing momentum equations (6 Reynolds stresses: $\rho \overline{u^{\prime 2}} ; \rho \overline{v^{\prime 2}} ; \rho \overline{w^{\prime 2}} ; \rho \overline{u^{\prime} v^{\prime}}$; $\left.\rho \overline{u^{\prime} w^{\prime}} ; \rho \overline{v^{\prime} w^{\prime}}\right)$ plus 3 additional unknowns per each scalar property $\varphi$ ( $\left(\overline{u^{\prime} \varphi^{\prime}} ; \overline{v^{\prime} \varphi^{\prime}} ; \overline{w^{\prime} \varphi^{\prime}}\right)$. The new system of governing time-averaged equations can become solvable again by introduction of appropriate turbulence model that will state and add differential equations, necessary to fulfil criterions of solvable system of equations. There exist 6 basic RANS turbulence models, varying in complexity and application areas: 1) Mixing length (+0eq), 2) Spalart-Allmaras (+1eq), 3) k- $(+2 e q)$,
4) $k-\omega$ ( $+2 e q$ ), 5) Algebraic stress (+2eq), 6) Reynolds stress ( +7 eq ) (Versteeg \& Malalasekera 2007, 66-98)

In order to understand the background of every listed turbulence model, one has to recall the Newton's law of viscosity (2.2.13) and introduce the Boussinesq approximation (2.3.25).

Using the suffix notation, Newton's law of viscosity can be rewritten in next form:

$$
\begin{equation*}
\tau_{i j}=\mu s_{i j}=\mu\left(\frac{\partial u_{i}}{\partial x_{j}}+\frac{\partial u_{j}}{\partial x_{i}}\right) \tag{2.3.24}
\end{equation*}
$$

The convection of suffix notation is that indices $i$ or $j=1$ correspond to the x -direction, $i$ or $j=2$ the y -direction and $i$ or $j=3$ the $z$-direction. For example:

$$
\tau_{12}=\tau_{x y}=\mu\left(\frac{\partial u_{1}}{\partial x_{2}}+\frac{\partial u_{2}}{\partial x_{1}}\right)=\mu\left(\frac{\partial u}{\partial y}+\frac{\partial v}{\partial x}\right)
$$

The experimental evidences show that turbulence decays unless there is shear in isothermal incompressible flows. Furthermore, turbulent stresses are found to increase as the mean rate of deformation increases. Those facts were used by Joseph Valentin Boussinesq, who proposed that Reynolds stresses are proportional to mean rates of deformation in next way:

$$
\begin{equation*}
\tau_{i j}=-\rho \overline{u_{\imath}{ }^{\prime} u_{\jmath}^{\prime}}=\mu_{t}\left(\frac{\partial U_{i}}{\partial x_{j}}+\frac{\partial U_{j}}{\partial x_{i}}\right)-\frac{2}{3} \rho k \delta_{i j} \tag{2.3.25}
\end{equation*}
$$

Where $k$ is the turbulent kinetic energy from formula 2.3.10, $\delta_{i j}$ is Kronecker delta ( $\delta_{i j}=1$ if $i=j$, and $\delta_{i j}=0$ if $i \neq j$ ) and $\mu_{t}$ is called turbulent or eddy viscosity.

By analogy, turbulent transport of a scalar is taken to be proportional to the gradient of mean value of the transported quantity:

$$
\begin{equation*}
-\rho \overline{u_{\imath}^{\prime} \varphi^{\prime}}=\Gamma_{t} \frac{\partial \Phi}{\partial x_{i}} \tag{2.3.26}
\end{equation*}
$$

Where $\Gamma_{t}$ is the turbulent or eddy diffusivity.
The relation between eddy viscosity and eddy exists and expressed by usage of Prandt//Schmidt number defined in next way:

$$
\begin{equation*}
\sigma_{t}=\frac{\mu_{t}}{\Gamma_{t}} \tag{2.3.27}
\end{equation*}
$$

Various flow experiments confirm that value $\sigma_{t}$ is constant, and hence most of free and commercial CFD software set the value $\sigma_{t}=1$. (Versteeg \& Malalasekera 2007, 68).
2.3.3 Mixing length model

The concept of mixing length model is based on next assumptions:

- Application is limited to 2D flows. Hence, the Reynolds stress $\tau_{x y}=\tau_{y x}$ is the only present turbulence shear stress component: $\tau_{x x}$ and $\tau_{y y}$ are assumed to be negligibly small.
- deduced from dimensional analysis:

$$
\begin{equation*}
\mu_{t}=C \rho v l \tag{2.3.28}
\end{equation*}
$$

Where:
$C$ - is a dimensionless constant of proportionality.
$\vartheta$ - is a turbulent velocity scale
$l$ - is turbulent length scale.
Turbulent velocity scale $\vartheta$ can be further expanded into:

$$
\begin{equation*}
\vartheta=c l\left|\frac{\partial U}{\partial y}\right| \tag{2.2.29}
\end{equation*}
$$

Where $c$ is non-dimensional constant.
Substitution of 2.3.29 to 2.3.28 and hiding of constants C and c into new length scale $l_{m}$ will yield into:

$$
\begin{equation*}
\mu_{t}=\rho l_{m}^{2}\left|\frac{\partial U}{\partial y}\right| \tag{2.3.30}
\end{equation*}
$$

The equation 2.3.30, combined with 2.3.25 and other previously stated assumption will result in so called Prandtl's mixing length model:

$$
\begin{equation*}
\tau_{x y}=\tau_{y x}=-\rho \overline{u^{\prime} v^{\prime}}=\rho l_{m}^{2}\left|\frac{\partial U}{\partial y}\right| \frac{\partial U}{\partial y} \tag{2.3.31}
\end{equation*}
$$

The same approach, applied to turbulent transport of arbitrary scalar quantity will yield into:

$$
\begin{equation*}
-\rho \overline{v^{\prime} \varphi^{\prime}}=\Gamma_{t} \frac{\partial \Phi}{\partial y} \tag{2.3.32}
\end{equation*}
$$

The Mixing length model finally allows to define unknown Reynolds stresses for 2D flows with no additional equations (Therefore it's also called 0 -equation model). The only things that must be the object consideration are values of $l_{m}$ and $\sigma_{t}$. The specification of these values can be found in book of H. K. Versteeg, W. Malalasekera "An Introduction to Computational Fluid Dynamics, The Finite Volume Method" second edition, pages 70 .

Advantages of the Mixing length model:

- easy and inexpensive implementation
- sufficiently accurate predictions for thin shear layers: jets, mixing layers, wakes and boundary layers
- well established


## Disadvantages:

- completely incapable of modelling flows with separation and recirculation
- completely incapable of describing flows with separation and recirculation
2.3.4 k- $\varepsilon$ model

The standard $\mathrm{k}-\varepsilon$ model provides an acceptable compromise between reliability, computational costs and accuracy, what makes $k-\varepsilon$ model to be apparently the most popular turbulence model, used in industry. This is a semi-empirical 2 -equation eddy-viscosity model, solving 2 additional equations for turbulent kinetic energy " $k$ " (2.3.10) and rate of energy dissipation per unit volume " $\varepsilon$ ".

To understand the concept of $k-\varepsilon$, one has to introduce the concept of the mean kinetic energy " $K$ " and the instantaneous kinetic energy " $k(t)$ ", that are defined in next way:

$$
\begin{equation*}
K=\frac{1}{2}\left(U^{2}+V^{2}+W^{2}\right) \tag{2.3.33}
\end{equation*}
$$

$$
\begin{equation*}
k(t)=K+k \tag{2.3.34}
\end{equation*}
$$

Another prerequisite for further model description is the decomposition of deformation rate tensor " $s_{i j}$ " to average and fluctuating part. Recalling formula 2.2.12, one can show that decomposition of " $s_{i j}$ " will hold as follows:

$$
\begin{equation*}
s_{i j}=S_{i j}+s^{\prime}{ }_{i j}=\frac{1}{2}\left[\frac{\partial U_{i}}{\partial x_{j}}+\frac{\partial U_{j}}{\partial x_{i}}\right]+\frac{1}{2}\left[\frac{\partial u^{\prime}{ }_{i}}{\partial x_{j}}+\frac{\partial u_{j}^{\prime}}{\partial x_{i}}\right] \tag{2.3.35}
\end{equation*}
$$

The scalar product of two tensors " $a_{i j}$ " and " $b_{i j}$ " is defined as follows:

$$
a_{i j} \cdot b_{i j}=a_{11} b_{11}+a_{12} b_{12}+a_{13} b_{13}+a_{21} b_{21}+a_{22} b_{22}+a_{23} b_{23}+a_{31} b_{31}+a_{32} b_{32}+a_{33} b_{33}
$$

It can be shown that governing equations for mean flow kinetic energy " $K$ " (2.3.36) and for turbulent kinetic energy " $k$ " (2.3.37) will take next form:

$$
\begin{equation*}
\frac{\partial(\rho K)}{\partial t}+\operatorname{div}(\rho K \mathbf{U})=\operatorname{div}\left(-P \mathbf{U}+2 \mu \mathbf{U} S_{i j}-\rho \mathbf{U} \overline{u^{\prime}{ }_{\imath} u_{\jmath}^{\prime}}\right)-2 \mu S_{i j} \cdot S_{i j}+\rho \overline{u_{\imath}^{\prime} u^{\prime}} . S_{i j} \tag{2.3.36}
\end{equation*}
$$

$$
\begin{equation*}
\frac{\partial(\rho k)}{\partial t}+\operatorname{div}(\rho k \mathbf{U})=\operatorname{div}\left(-\overline{p^{\prime} \mathbf{u}^{\prime}}+2 \mu \overline{\mathbf{u}^{\prime} s^{\prime}{ }_{l j}}-\rho \frac{1}{2} \overline{u^{\prime} \cdot u^{\prime} \cdot{ }_{\imath} u^{\prime}{ }_{\jmath}}\right)-2 \mu \overline{{s^{\prime}}^{\prime} \cdot s^{\prime}{ }_{l \jmath}}+\rho \overline{u^{\prime}{ }_{\imath} u^{\prime}} \cdot S_{i j} \tag{2.3.37}
\end{equation*}
$$

The second term on RHS of 2.3.37 is usually written as product of density " $\rho$ " and the rate of dissipation of turbulent kinetic energy per unit mass " $\varepsilon$ ". Therefore " $\varepsilon$ " is defined as follows:

$$
\begin{equation*}
\varepsilon=2 \frac{\mu}{\rho} \overline{s^{\prime}{ }_{l \jmath} \cdot s^{\prime}{ }_{l \jmath}} \tag{2.3.38}
\end{equation*}
$$

It is also possible to derive the exact differential governing equation for " $\varepsilon$ ", but it contains many unknowns, and hence the standard $k-\varepsilon$ model is based on next assumptions for velocity scale " $\vartheta$ " and length scale " $l$ ":

$$
\begin{equation*}
\vartheta=k^{1 / 2} \tag{2.3.39}
\end{equation*}
$$

$$
\begin{equation*}
l=\frac{k^{3 / 2}}{\varepsilon} \tag{2.3.40}
\end{equation*}
$$

The substitution of 2.3.39 and 2.3.40 to formula of eddy viscosity will give next relation:

$$
\begin{equation*}
\mu_{t}=C \rho \vartheta l=\rho C_{\mu} \frac{k^{2}}{\varepsilon} \tag{2.3.41}
\end{equation*}
$$

Where " $C_{\mu}$ " - is a dimensionless constant.
The formula 2.3.41 itself is an assumption of isotropic eddy viscosity, allowing to state two transport equations of standard $k-\varepsilon$ model:

$$
\begin{equation*}
\frac{\partial(\rho k)}{\partial t}+\operatorname{div}(\rho k \mathbf{U})=\operatorname{div}\left[\frac{\mu_{t}}{\sigma_{k}} \operatorname{grad}(k)\right]+2 \mu_{t} S_{i j} . S_{i j}-\rho \varepsilon \tag{2.3.42}
\end{equation*}
$$

$$
\begin{equation*}
\frac{\partial(\rho \varepsilon)}{\partial t}+\operatorname{div}(\rho \varepsilon \mathbf{U})=\operatorname{div}\left[\frac{\mu_{t}}{\sigma_{\varepsilon}} \operatorname{grad}(\varepsilon)\right]+C_{1 \varepsilon} \frac{\varepsilon}{k} 2 \mu_{t} S_{i j} . S_{i j}-C_{2 \varepsilon} \frac{\varepsilon^{2}}{k} \tag{2.3.43}
\end{equation*}
$$

Where $\quad C_{\mu}=0.09 ; \sigma_{k}=1 ; \sigma_{\varepsilon}=1.3 ; C_{1 \varepsilon}=1.44 ; C_{2 \varepsilon}=1.92$ - are empirically defined dimensionless constants, suitable for wide range of flows.
The Reynolds stresses are found, using following Boussinesq approximation:

$$
-\rho \overline{u^{\prime}{ }_{\imath} u^{\prime}{ }_{J}}=\mu_{t}\left(\frac{\partial U_{i}}{\partial x_{j}}+\frac{\partial U_{j}}{\partial x_{i}}\right)-\frac{2}{3} \rho k \delta_{i j}=2 \mu_{t} S_{i j}-\frac{2}{3} \rho k \delta_{i j}
$$

In order to run $\mathrm{k}-\varepsilon$ model appropriately, some CFD codes in addition to turbulence intensity might also ask to specify the values of " $k$ " and " $\varepsilon$ " for system inlet. This can be done either by reviewing literature, covering particular cases of study, which is more preferable option, or using next formulas, connecting " $k$ " and " $\varepsilon$ " with turbulence intensity " $T_{i}$ " and length scale " $l$ ":

$$
\begin{equation*}
k=\frac{2}{3}\left(U_{r e f} T_{i}\right)^{2} \tag{2.3.44}
\end{equation*}
$$

$$
\begin{equation*}
\varepsilon=C_{\mu}^{3 / 4} \frac{k^{3 / 2}}{l} \tag{2.3.45}
\end{equation*}
$$

$$
\begin{equation*}
l=0.07 L \tag{2.3.46}
\end{equation*}
$$

Where " $L$ " - is a characteristic length of equipment (equivalent pipe diameter). (Versteeg \& Malalasekera 2007, 72-88).

Equivalent pipe diameter on-line calculator with some explanation theory:
http://www.engineeringtoolbox.com/equivalent-diameter-d_205.html In cases, when equivalent pipe diameter is not obvious to define it's sufficient either to use default settings in CFD code (if exist) or arbitrary finite and small values for " $\varepsilon$ ". (Versteeg \& Malalasekera 2007, 77).

In addition to standard $\mathrm{k}-\varepsilon$ (SKE) most of free and commercial CFD codes provide users with two more advanced variants of $k-\varepsilon$ model: $k-\varepsilon$ RNG (Renormalization Groups) and $k-\varepsilon$ RKE (Realizable).
k- $\varepsilon$ RNG model instead of using empirically defined constants $C_{\mu} ; \sigma_{k} ; \sigma_{\varepsilon} ; C_{1 \varepsilon} ; C_{2 \varepsilon}$ resolves them using statistical methods, making it more precise for wider range of more complex flows.
$\mathrm{k}-\varepsilon$ RKE model is an improvement of standard $\mathrm{k}-\varepsilon$ model, varying in next points:

- $k-\varepsilon$ RKE contains a new formulation for the turbulent viscosity with varying parameter " $C_{\mu}$ " that was assumed to be constant for standard $\mathrm{k}-\varepsilon$ model.
- A new transport equation for the dissipation rate " $\varepsilon$ " is derived from exact transport equation of the mean-square vorticity fluctuation.

Unlike $k-\varepsilon$ SKE or $k-\varepsilon$ RNG, $k-\varepsilon$ RKE model satisfies several constraints of physics of turbulent flows, making RKE potentially the most accurate variation of $k-\varepsilon$ model. Note that all $k-\varepsilon$ models variation are preferred to use only for fully turbulent flows (high Reynolds numbers), due to fully turbulent flow assumption as a basement of the whole model. In addition to $k-\varepsilon$, there also exist other two-equation models such as Wilcox $k-\omega$, Menter SST $k-\omega$, algebraic stress equation model and non-linear $k-\varepsilon$. (Versteeg \& Malalasekera 2007, 90-95).

Unlike $k-\varepsilon$ SKE, $k-\varepsilon$ RNG or $k-\varepsilon$ RKE, Wilcox $k-\omega$ shows the best performance for flows with low Reynolds numbers and very accurate results for flows in near-wall regions. It's success in near-wall computations even yielded to creation of hybrid Menter SST k- $\omega$ model,
which solves the near-wall flow (boundary layer regions) using Wilcox $k$ $\omega$, and $k-\varepsilon$ for the rest of the flow regions. This makes Menter SST $k-\omega$ model to be a perfect option for general cases of subsonic flows passing various geometric objects (obstacles). (Versteeg \& Malalasekera 2007, 91 -93).

So far, all covered turbulence models use the assumption of isotropic eddy viscosity, which isn't truth for reality. In cases, when anisotropy of turbulence is important to encounter but computational powers limit the usage to only two-equation and simpler models, one can use the algebraic stress model. However, this modification of $k-\varepsilon$ model suffers reliability and re-quires experimental validation before results of usage of this model influence the further design work. At last, algebraic stress model is currently being overshadowed by non-linear $k-\varepsilon$ models, that able to show results close to more complicated and precise sevenequation models but still using only two extra transport equations. (Versteeg \& Malalasekera 2007, 93-95).

### 2.3.5 Spalart-Allmaras turbulence model

Spalart-Allmaras is one-equation turbulence model based on transport of kinematic eddy viscosity parameter " $\tilde{v}$ ", which is connected with dynamic eddy viscosity " $\mu_{t}$ " by next relation:

$$
\begin{equation*}
\mu_{t}=\rho \tilde{v} f_{v 1} \tag{2.3.47}
\end{equation*}
$$

Where " $f_{v 1}$ " is so-called wall-damping function $f_{v 1}=f_{v 1}\left(\frac{\tilde{v}}{v}\right) ; \quad\left(v=\frac{\mu}{\rho}\right)$, which behaves in such way, that it tends to unity as Reynolds number increases and tends to zero at wall boundaries:

$$
\begin{equation*}
f_{v 1}=\frac{\left(\frac{\tilde{v}}{v}\right)^{3}}{\left(\frac{\tilde{v}}{v}\right)^{3}-C_{v 1}^{3}} \tag{2.3.48}
\end{equation*}
$$

Where $C_{\nu 1}=7.1$
Substitution of 2.3.47 to 2.3.25 and neglecting of term " $\frac{2}{3} \rho k \delta_{i j}$ " will give a formula for Reynolds stresses:

$$
\begin{equation*}
\tau_{i j}=-\rho \overline{u_{\imath}^{\prime}{ }_{\imath} u^{\prime}}=2 \mu_{t} S_{i j}=\rho \tilde{v} f_{v 1}\left(\frac{\partial U_{i}}{\partial x_{j}}+\frac{\partial U_{j}}{\partial x_{i}}\right) \tag{2.3.49}
\end{equation*}
$$

The transport equation for " $\tilde{v}^{\prime \prime}$ holds as follows:

$$
\begin{align*}
& \frac{\partial(\rho \tilde{v})}{\partial t}+\operatorname{div}(\rho \tilde{v} \mathbf{U}) \\
&=\frac{1}{\sigma_{v}} \operatorname{div}\left[(\mu+\rho \tilde{v}) \operatorname{grad}(\tilde{v})+C_{b 2} \rho \frac{\partial \tilde{v}}{\partial x_{k}} \frac{\partial \tilde{v}}{\partial x_{k}}\right]+C_{b 1} \rho \tilde{v} \widetilde{\Omega} \\
&-C_{w 1} \rho\left(\frac{\tilde{v}}{\kappa y}\right)^{2} f_{w} \tag{2.3.50}
\end{align*}
$$

Where:

$$
\begin{equation*}
\widetilde{\Omega}=\Omega+\frac{\tilde{v}}{(\kappa y)^{2}} f_{v 2} \tag{2.3.51}
\end{equation*}
$$

$$
\Omega=\sqrt{2 \Omega_{i j} \Omega_{i j}}
$$

(2.3.52) (mean vorticity)

$$
\begin{equation*}
\Omega_{i j}=\frac{1}{2}\left(\frac{\partial U_{i}}{\partial x_{j}}-\frac{\partial U_{j}}{\partial x_{i}}\right) \tag{2.3.53}
\end{equation*}
$$

$$
f_{v 2}=1-\frac{\frac{\tilde{v}}{v}}{1+\frac{v}{v} f_{v 1}}
$$

$$
f_{w}=g\left[\frac{1+C_{w 3}^{6}}{g^{6}+C_{w 3}^{6}}\right]^{\frac{1}{6}}
$$

$$
g=r+C_{w 2}\left(r^{6}-r\right)
$$

$$
\begin{equation*}
r=\min \left[\frac{\tilde{v}}{\widetilde{\Omega} \kappa^{2} y^{2}} ; 10\right] \tag{2.3.55}
\end{equation*}
$$

And " $y$ " is the smallest distance to solid wall (same as " $y$ " in 2.3.16).

Table 3. Constants used in Spalart-Allmaras model.

| Constant | Value |
| :---: | :---: |
| $\sigma_{v}$ | $2 / 3$ |
| $\kappa$ | 0.4187 |
| $C_{b 1}$ | 0.1355 |
| $C_{b 2}$ | 0.622 |
| $C_{w 1}$ | $C_{b 1}+\kappa^{2} \frac{1+C_{b 2}}{\sigma_{v}}$ |
| $C_{w 2}$ | 0.3 |
| $C_{w 3}$ | 2 |

By default, Spalart-Allmaras requires user to model the grid to be fine enough to capture the viscous sublayer, using the law of the wall. However, some commercial codes like ANSYS FLUENT provide users with Enhanced Wall Treatment ( $y^{+}$insensitive wall treatment), which allows the application of the model independent of the near-wall grid resolution. Finally, Spalart-Allmaras turbulence model is the model, specifically designed to model the problems of aerodynamics, what ensures the highest accuracy for external aerodynamic flows on the one hand, but suffers from inaccuracy in the rest Fluid Dynamics' fields on the other hand (https://www.cfd-online.com/Wiki/Spalart-Allmaras_model). (Versteeg \& Malalasekera 2007, 89-90).
2.3.6 Reynolds stress equation models (RSM)

Reynolds stress model (RSM) is the seven-equation model, giving extra transport equations to every unique from six components of stress tensor " $\tau_{i j}$ " and one more equation for either " $\varepsilon$ " or " $\omega$ ".

The transport equations for Reynolds stresses however are described by variable called kinematic Reynolds stresses:

$$
\begin{equation*}
R_{i j}=-\frac{\tau_{i j}}{\rho}=\overline{u_{\imath}^{\prime} u_{J}^{\prime}} \tag{2.3.56}
\end{equation*}
$$

The exact transport equation for " $R_{i j}$ " takes next form:

$$
\begin{equation*}
\frac{\partial\left(\rho R_{i j}\right)}{\partial t}+\operatorname{div}\left(\rho R_{i j} \mathbf{U}\right)=P_{i j}+D_{i j}-\varepsilon_{i j}+\Pi_{i j}+\Omega_{i j} \tag{2.3.57}
\end{equation*}
$$

Where:

$$
\begin{gather*}
P_{i j}=-\left(R_{i m} \frac{\partial U_{j}}{\partial x_{m}}+R_{j m} \frac{\partial U_{i}}{\partial x_{m}}\right)  \tag{2.3.58}\\
D_{i j}=\operatorname{div}\left(\frac{v_{t}}{\sigma_{k}} \operatorname{grad}\left(R_{i j}\right)\right) \\
v_{t}=C_{\mu} \frac{k^{2}}{\varepsilon} \\
C_{\mu}=0.09 ; \sigma_{k}=1  \tag{2.3.59}\\
\varepsilon_{i j}=\frac{2}{3} \varepsilon \delta_{i j} \tag{2.3.60}
\end{gather*}
$$

$$
\Pi_{i j}=-C_{1} \frac{\varepsilon}{k}\left(R_{i j}-\frac{2}{3} k \delta_{i j}\right)-C_{2}\left(P_{i j}-\frac{2}{3} P \delta_{i j}\right)
$$

$$
\begin{equation*}
C_{1}=1.8 ; C_{2}=0.6 \tag{2.3.61}
\end{equation*}
$$

$$
\begin{equation*}
\Omega_{i j}=-2 \omega_{k}\left(R_{j m} e_{i k m}+R_{i m} e_{j k m}\right) \tag{2.3.62}
\end{equation*}
$$

Where:
$\omega_{k}$ - is the rotation vector, $e_{i j k}=1$ if $i, j$ and $k$ are different and in cyclic order, $e_{i j k}=-1$ if $i, j$ and $k$ are different and in anti-cyclic order; and $e_{i j k}=0$ if any two indices are same.

The transport equation for scalar dissipation rate " $\varepsilon$ " for sake of simplicity is same as for standard $k-\varepsilon$ model (2.3.43). This is valid for most of commercial CFD codes, in spite the fact that more precisely defined equation exist. (Versteeg \& Malalasekera 2007, 80-84).

More detailed information about turbulence models can be found via following links:

- https://turbmodels.larc.nasa.gov/index.html
- https://www.cfd-online.com/Wiki/Turbulence modeling


### 2.3.7 Direct Numerical Simulation (DNS) and Large Eddy Simulation (LES)

In Direct Numerical Simulation system of governing transient equations 2.2.1, 2.2.20, 2.2.21 and 2.2.22 solved directly without implementation of any turbulence model and Reynolds averaging at all (not a RANS model). As a consequence, DNS demands extremely fine mesh an sufficiently small time steps in order of simulate the motion of eddies with smallest size and highest rotational frequencies. Therefore, DNS demands computational powers that can be only satisfied by modern supercomputers, making this method unsuitable for commercial usage. And even usage of supercomputers so far haven't allowed to use this method on complex geometries with highly turbulent flows. Finally, at the moment DNS can be applied only to incompressible, simple-geometry and low-Reynolds-number flows.

In spite such limitations, scientists came up with computational technique called Large Eddy Simulation (LES), that can be considered as simplification of DNS applicable for conventional computers, allowing to use coarser meshing (not a RANS model too).

The ideas behind LES are two empirically proven facts:

- Most of kinetic turbulent energy is contained in largest eddies in flow, meaning that smaller eddies play relatively negligible role in turbulence effects.
- Anisotropy of eddies increases as the size of eddy increases.

Both of these facts are used to exclude the modelling of small eddies from simulation running process (that are later resolved using so-called sub-grid-scale (SGS) model), significantly reducing computational costs. The exclusion of small eddies is done via spatial filtering operation:

$$
\begin{equation*}
\bar{\varphi}(\mathbf{x}, t)=\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} G\left(\mathbf{x}, \mathbf{x}^{\prime}, \Delta\right) \varphi\left(\mathbf{x}^{\prime}, t\right) d x_{1} d x_{2} d x_{3} \tag{2.3.63}
\end{equation*}
$$

Where:
$\bar{\varphi}(\mathbf{x}, t)$ - is a filtered function (in this subsection overbar indicates filtering, not averaging);
$\varphi(\mathbf{x}, t)$ - is original (unfiltered) function;
$\Delta$ - is a filter cutoff width;
$G\left(\mathbf{x}, \mathbf{x}^{\prime}, \Delta\right)$ - is a filter function.
There exist 3 filtering functions that were successfully implemented to CFD: Top-hat (box) filter, Gaussian filter and Spectral cutoff. However, only Top-hat filter so far is widely used in commercial and free software,
while rest two can be found only in research literature. The box filter is defined as follows:

$$
G\left(\mathbf{x}, \mathbf{x}^{\prime}, \Delta\right)=\left\{\begin{array}{c}
\frac{1}{\Delta^{3}} ;\left|\mathbf{x}-\mathbf{x}^{\prime}\right| \leq \frac{\Delta}{2}  \tag{2.3.64}\\
0 ;\left|\mathbf{x}-\mathbf{x}^{\prime}\right|>\frac{\Delta}{2}
\end{array}\right.
$$

Cutoff width " $\Delta$ " is defined in next way:

$$
\begin{equation*}
\Delta=\sqrt[3]{\Delta x \Delta y \Delta z} \tag{2.3.65}
\end{equation*}
$$

Where $\Delta x \Delta y \Delta z$ - is the volume of biggest (by volume) cell in grid.
Even though it is possible to set the cutoff width to be even smaller, there is no point in doing that because the grid resolution simply will not be able to model them adequately. Besides, cutoff width reduction will also negatively affect sub-grid-scale (SGS) stresses - the stresses, resolved from rejected smaller eddies. Increasing of cutoff width, on the other hand, will result to lesser amount of modelled eddies, what yields to lower accuracy of simulation. Therefore, it is the best to keep formula 2.3.65 as close as possible.

Similarly to RANS models, LES requires appropriate boundary conditions, in order to obtain a well-posed problem. The boundary conditions are adjusted similarly to RANS models. However, some difference exists:

- Adjustment of dimensionless constant " $C_{S G S}$ ". The value of this constant varies from one to other SGS model and usually lay in range from 0.1 to 0.24 . The best way to determine the exact value of this constant is revising research literature with similar cases.
- Cutoff width setup (see 2.3.65)
- First wall mesh cell height must be computed strictly for $y^{+} \leq 1$, independently on simplicity of geometry. Usage of wall functions is also possible.
- Steady flow is solved using transient governing equation, what means that several time steps must get passed, before final state of flow is formed. The best way to determine time gap from initial to final state of flow is revision of research literature.
- Inflow conditions are the most challenging part of setup.There exist several approaches to set correct conditions:

1. Preparatory transient simulation of same problem but using RSM model. Obtained Reynolds stresses at the inlet are further transferred through corresponding autocorrelation and cross-correlation functions, that form random Gaussian
perturbations, to LES
2. Extension of computational domain. Long upstream distances are required in order to ensure the generation of fully developed flow from turbulence-free reservoirs
3. Direct specification of shear stresses and velocity profiles in inlet. (Versteeg \& Malalasekera 2007, 98-114).

### 2.4 Finite volume method and solution schemes

### 2.4.1 Finite volume method for diffusion problems

Pure diffusion problems are one of the easiest problems that finite volume method can handle. In physics and engineering pure diffusion problems mostly involve problems of heat transfer in solids. In spite the fact that pure diffusion model is not sufficient for fluids due to involvement of convection, discretization procedure for solids and fluids is similar. There-fore in this section problems of pure steady-state diffusion will be covered first, and convection-diffusion problems with pressure-velocity coupling later. Pure diffusion equation for threedimensional problems is written as follows:

$$
\begin{equation*}
\operatorname{div}(\Gamma \operatorname{grad}(\varphi))+S_{\varphi}=0 \tag{2.4.1}
\end{equation*}
$$

Integration over control volume and application of Gauss's divergence theorem will result to next form of 2.4.1:

$$
\begin{align*}
& \int_{C V} \operatorname{div}(\Gamma \operatorname{grad}(\varphi)) d V+\int_{C V} S_{\varphi} d V= \\
& \int_{C S} \mathbf{n} \cdot(\Gamma \operatorname{grad}(\varphi)) d A+\int_{C V} S_{\varphi} d V=0 \tag{2.4.2}
\end{align*}
$$

Simplification of 2.4.2 to one-dimensional case will result to following equation:

$$
\begin{equation*}
\int_{C V} \frac{d}{d x}\left(\Gamma \frac{d \varphi}{d x}\right) d V+\int_{C V} S_{\varphi} d V=0 \tag{2.4.3}
\end{equation*}
$$

Similarly to other numerical methods, Finite volume method involves discretization. Uniquely for FVM, discretization is applied to control volume, resulting to its partition to finite amount of nodal points, surrounded by smaller control volumes (cells). Each nodal point corresponds to unique set of linearized governing equations specific to this smaller control volume and coupled with governing equations of surrounding cells at certain degree. (Versteeg \& Malalasekera 2007, 115).

The example of one-dimensional discretization is illustrated in Figure 10. The continuous line $A B$ is divided to five control volumes with nodal
points placed in centre of cell, and values of property $\phi$ at boundaries " $A$ " and " $B$ " are specified.


Figure 10. Spatial discretization of one-dimensional space. (Versteeg \& Malalasekera 2007, 116).

In order to make the specification of parameters for every single cell of control volume to be convenient, the convention, illustrated in Figure 11, will be used further.


Figure 11. Notation system, used for one-dimensional CFD applications. (Versteeg \& Malalasekera 2007, 116).

Where:
P - is the index used to mark the properties of flow at target node;
W - is the index used to mark the properties of flow at neighbour node, at west (on the left) from target node;
E - is the index used to mark the properties of flow at neighbour cell, at east (on the right) from target node;
w - is the index used to mark the properties of flow at west side control volume face of target node;
e - is the index used to mark the properties of flow at east side control volume face ;

Using the assumption of linear change from one nodal point to another, one can show that equation 2.4 .3 for particular point " P " will transform to next form:

$$
\begin{equation*}
\int_{C V} \frac{d}{d x}\left(\Gamma \frac{d \varphi}{d x}\right) d V+\int_{C V} S_{\varphi} d V=\left(\Gamma A \frac{d \varphi}{d x}\right)_{e}-\left(\Gamma A \frac{d \varphi}{d x}\right)_{w}+\bar{S} \Delta V=0 \tag{2.4.4}
\end{equation*}
$$

Where:
$A$ - is a cross-sectional area of cell face;
$\Delta V$ - is a volume of target cell;
$\bar{S}$ - is averaged value of S over control volume.
In order to determine the diffusion coefficients " $\Gamma$ " at west and east side faces of cell, one again can use linear interpolation with respect to diffusion coefficients at neighbouring nodal points:

$$
\begin{gather*}
\Gamma_{w}=\frac{\Gamma_{W}-\Gamma_{P}}{2} \\
\Gamma_{e}=\frac{\Gamma_{P}-\Gamma_{E}}{2} \tag{2.4.5}
\end{gather*}
$$

Diffusion terms can be linearized as follows:

$$
\begin{gather*}
\left(\Gamma A \frac{d \varphi}{d x}\right)_{w}=\Gamma_{w} A_{w}\left(\frac{\varphi_{P}-\varphi_{W}}{\delta x_{W P}}\right) \\
\left(\Gamma A \frac{d \varphi}{d x}\right)_{e}=\Gamma_{e} A_{e}\left(\frac{\varphi_{E}-\varphi_{P}}{\delta x_{P E}}\right) \tag{2.4.6}
\end{gather*}
$$

For source term $\bar{S} \Delta V$ linearization applied as follows:

$$
\begin{equation*}
\bar{S} \Delta V=S_{u}+S_{P} \varphi_{P} \tag{2.4.7}
\end{equation*}
$$

Substitution of 2.4.6 and 2.4.7 to 2.4.4 will results to:

$$
\begin{equation*}
\Gamma_{e} A_{e}\left(\frac{\varphi_{E}-\varphi_{P}}{\delta x_{P E}}\right)-\Gamma_{w} A_{w}\left(\frac{\varphi_{P}-\varphi_{W}}{\delta x_{W P}}\right)+\left(S_{u}+S_{P} \varphi_{P}\right)=0 \tag{2.4.8}
\end{equation*}
$$

Which can be further rearranged to:

$$
\begin{equation*}
\left(\frac{\Gamma_{e}}{\delta x_{P E}} A_{e}+\frac{\Gamma_{w}}{\delta x_{W P}} A_{w}-S_{P}\right) \varphi_{P}=\left(\frac{\Gamma_{w}}{\delta x_{W P}} A_{w}\right) \varphi_{W}+\left(\frac{\Gamma_{e}}{\delta x_{P E}} A_{e}\right) \varphi_{E}+S_{u} \tag{2.4.9}
\end{equation*}
$$

The resulting equation 2.4 .9 is a simple algebraic equation allowing to solve value " $\varphi$ " for particular nodal point $P$. By refinement of the grid (adding the nodal points) one can solve " $\varphi$ " with higher accuracy (" $\varphi$ " appears to be more specifically determined in control volume). On the other hand, mesh refinement yields to increased amount of algebraic equations (1 algebraic equation per nodal point), what results to increase size of corresponding matrix of system of algebraic equations, finally leading to in-creased computational time. The outcome of 2.4.9 can be further generalized to next form:

$$
\begin{equation*}
a_{P} \varphi_{P}=a_{W} \varphi_{W}+a_{E} \varphi_{E}+S_{u} \tag{2.4.10}
\end{equation*}
$$

With corresponding table for coefficients " a ":

Table 4. Coefficients " a " for 1D diffusion problems.

| $a_{W}$ | $a_{E}$ | $a_{P}$ |
| :---: | :---: | :---: |
| $\frac{\Gamma_{w} A_{w}}{\delta x_{W P}}$ | $\frac{\Gamma_{e} A_{e}}{\delta x_{P E}}$ | $a_{W}+a_{E}+S_{p}$ |

2.4.9 or 2.4.10 together with Table 4 form the mathematical model, sufficient to solve one-dimensional problems. However 1D diffusion models are rarely used in real engineering work due to insufficient accuracy especially in complex geometries. Therefore one has to extend existing methods for one-dimensional cases to two- and threedimensional cases by adding extra terms to 2.4.10:

$$
\begin{equation*}
a_{P} \varphi_{P}=a_{W} \varphi_{W}+a_{E} \varphi_{E}+a_{S} \varphi_{S}+a_{N} \varphi_{N}+S_{u} \tag{2.4.11}
\end{equation*}
$$

Where indices " S " and " N " stand for southern and northern neighbouring nodal points according to Figure 12.


Figure 12. Notation used for two-dimensional CFD applications. (Versteeg \& Malalasekera 2007, 129).

Corresponding "a" coefficients are summarized in "Table 5".

Table 5. Coefficients " a " for 2D diffusion problems.

| $a_{W}$ | $a_{E}$ | $a_{S}$ | $a_{N}$ | $a_{P}$ |
| :---: | :---: | :---: | :---: | :---: |
| $\frac{\Gamma_{w} A_{w}}{\delta x_{W P}}$ | $\frac{\Gamma_{e} A_{e}}{\delta x_{P E}}$ | $\frac{\Gamma_{S} A_{S}}{\delta y_{S P}}$ | $\frac{\Gamma_{n} A_{n}}{\delta y_{P N}}$ | $a_{W}+a_{E}+a_{S}+a_{N}+S_{p}$ |

Where:
$A_{w}=A_{e}=\Delta y ;$ And $A_{s}=A_{n}=\Delta x$.

Applying same principles, one can show that equation 2.4.11, once expanded to three-dimensional problems, will transform to next form:

$$
a_{P} \varphi_{P}=a_{W} \varphi_{W}+a_{E} \varphi_{E}+a_{S} \varphi_{S}+a_{N} \varphi_{N}+a_{B} \varphi_{B}+a_{T} \varphi_{T}+S_{u}
$$

Where indices " T " and " B " stand for "Top" and "Bottom" as illustrated in Figure 13.


Figure 13. Notation used for two-dimensional CFD applications. (Versteeg \& Malalasekera 2007, 131).

Corresponding "a" coefficients for three-dimensional cases are summarized in "Table 6".

Table 6. Coefficients "a" for 3D diffusion problems.

| $a_{W}$ | $\frac{\Gamma_{w} A_{w}}{\delta x_{W P}}$ |
| :---: | :---: |
| $a_{E}$ | $\frac{\Gamma_{e} A_{e}}{\delta x_{P E}}$ |
| $a_{S}$ | $\frac{\Gamma_{s} A_{s}}{\delta y_{S P}}$ |
| $a_{N}$ | $\frac{\Gamma_{n} A_{n}}{\delta y_{P N}}$ |
| $a_{B}$ | $\frac{\Gamma_{b} A_{b}}{\delta z_{B P}}$ |
| $a_{T}$ | $\frac{\Gamma_{t} A_{t}}{\delta z_{P T}}$ |
| $a_{P}$ | $a_{W}+a_{E}+a_{S}+a_{N}+a_{B}+a_{T}-S_{P}$ |

Observing the pattern in 2.4.10, 2.4.11 and 2.4.12, one can write generalized form of these equations in next way:

$$
\begin{equation*}
a_{P} \varphi_{P}=\sum a_{n b} \varphi_{n b}+S_{u} \tag{2.4.13}
\end{equation*}
$$

And,

$$
\begin{equation*}
a_{P}=\sum a_{n b}-S_{P} \tag{2.4.14}
\end{equation*}
$$

Where index "nb" corresponds to neighbouring nodes relatively to target node. (Versteeg \& Malalasekera 2007, 115-133).

### 2.4.2 Convection-diffusion problems

General form of convection-equation is read as follows:

$$
\begin{equation*}
\int_{A} \mathbf{n} \cdot(\rho \varphi \mathbf{u}) d A=\int_{A} \mathbf{n} \cdot\left(\Gamma(\operatorname{grad}(\varphi)) d A+\int_{C V} S_{\varphi} d V\right. \tag{2.4.15}
\end{equation*}
$$

Simplification of 2.4.15 to one-dimensional and neglecting of source term case will result to:

$$
\begin{equation*}
\frac{d(\rho u \varphi)}{d x}=\frac{d}{d x}\left(\Gamma\left(\frac{d \varphi}{d x}\right)\right) \tag{2.4.16}
\end{equation*}
$$

In addition to 2.4.16, one also has to consider the continuity equation:

$$
\begin{equation*}
\frac{d(\rho u)}{d x}=0 \tag{2.4.17}
\end{equation*}
$$

To apply discretization to 2.4.16 and 2.4.17, the notation similar to diffusion problems, shown in Figure 14, is used with only difference of presence of velocity components " $u$ ", " $v$ " and " $w$ ".


Figure 14. Notation, used for one-dimensional convection-diffusion problems. (Versteeg \& Malalasekera 2007, 135).

With given convention, one can transform equations 2.4.16 and 2.4.17 in linearized form:

$$
\begin{equation*}
(\rho u A \varphi)_{e}-(\rho u A \varphi)_{w}=\left(\Gamma A \frac{d \varphi}{d x}\right)_{e}-\left(\Gamma A \frac{d \varphi}{d x}\right)_{w} \tag{2.4.18}
\end{equation*}
$$

For transport equation of " $\varphi$ ", and for continuity equation:

$$
\begin{equation*}
(\rho u A)_{e}-(\rho u A)_{w}=0 \tag{2.4.19}
\end{equation*}
$$

For the sake of convenience, one can introduce next replacements for 2.4.18 and 2.4.19:

$$
\begin{equation*}
F=\rho u \tag{2.4.20}
\end{equation*}
$$

And,

$$
\begin{equation*}
D=\frac{\Gamma}{\delta x} \tag{2.4.21}
\end{equation*}
$$

Particularly for 2.4.18 and 2.4.19, one can get next coefficients:

$$
\begin{aligned}
& F_{w}=(\rho u)_{w} \\
& F_{e}=(\rho u)_{e} \\
& D_{w}=\frac{\Gamma_{w}}{\delta x_{W P}}
\end{aligned}
$$

$$
\begin{equation*}
D_{e}=\frac{\Gamma_{e}}{\delta x_{P E}} \tag{2.4.22}
\end{equation*}
$$

Substitution of 2.4.22 to 2.4.18 and to 2.4.19 and assumption that $A_{w}=A_{e}=A$ will give next equations:

$$
\begin{equation*}
F_{e} \varphi_{e}-F_{w} \varphi_{w}=D_{e}\left(\varphi_{E}-\varphi_{P}\right)-D_{w}\left(\varphi_{P}-\varphi_{W}\right) \tag{2.4.23}
\end{equation*}
$$

$$
\begin{equation*}
F_{e}-F_{w}=0 \tag{2.4.24}
\end{equation*}
$$

There exists several schemes to reduce number of variables in 2.4 .23 by representing values " $\varphi$ " at faces (" $\varphi$ " with lower-case letter indices) as sum of nodal terms (" $\varphi$ " with capital letter indices). However, the simplest scheme is called "The central differencing scheme", which represents " $\varphi$ " at faces as follows:

$$
\begin{align*}
& \varphi_{e}=\frac{\varphi_{P}+\varphi_{E}}{2} \\
& \varphi_{w}=\frac{\varphi_{W}+\varphi_{P}}{2} \tag{2.4.25}
\end{align*}
$$

Substitution of 2.4.25 to 2.4.23 will result to next equation:

$$
\begin{equation*}
a_{P} \varphi_{P}=a_{W} \varphi_{W}+a_{E} \varphi_{E} \tag{2.4.26}
\end{equation*}
$$

Table 7. "a" coefficients for one-dimensional convection-diffusion problems applying central differencing scheme.

| $a_{W}$ | $a_{E}$ | $a_{P}$ |
| :---: | :---: | :---: |
| $D_{w}+\frac{F_{w}}{2}$ | $D_{e}-\frac{F_{e}}{2}$ | $a_{W}+a_{E}+\left(F_{e}-F_{w}\right)$ |

In addition to central differencing schemes, several other discretization schemes that might perform better than central differencing scheme, but might give unrealistic result. In order to avoid such problem, one must consider the next fundamental properties of any discretization scheme:

- Conservativeness
- Boundedness
- Transportiveness

Conservativness is the property, telling how well properties " $\phi$ " obey conservativeness in overall bounded control volume. Discretization schemes, that satisfy conservatives are also called consistent. The examples of consistent schemes are Central Differencing Scheme and QUICK. The example of inconsistent scheme, using quadratic interpolation is shown in Figure 15 below.


Figure 15. Inconsistent scheme, using quadratic interpolation. (Versteeg \& Malalasekera 2007, 142)

In this example gradients of flux " $\phi$ " at west face of cell " 2 " and east face of cell " 3 " can be different, what yields to unsatisfied conservation condition. (Versteeg \& Malalasekera 2007, 134-143).

Boundedness is the property of scheme telling if the solution will converge after several iterations, that are necessary in CFD to handle nonlinear (product of unknown functions) terms of governing equations. The sufficient condition for convergence of iterative method is described by Scar-borough criterion:

$$
\frac{\sum\left|a_{n b}\right|}{\left|a_{P}^{\prime}\right|}\left\{\begin{array}{c}
\leq 1 \text { at all nodes }  \tag{2.4.27}\\
<1 \text { at one node at least }
\end{array}\right.
$$

Where $a_{P}^{\prime}$ denotes the coefficient at target nodal point and $a_{n b}$ denotes coefficients of all neighbouring nodes involved in linear algebraic equation of " $\varphi_{P}$ ". This criterion can be satisfied by appropriate adjustment of source term coefficients " S ". (Versteeg \& Malalasekera 2007, 143).

Transportiveness is the property characterized by Peclet number:

$$
\begin{equation*}
P e=\frac{F}{D}=\frac{\rho u}{\Gamma / \delta x} \tag{2.4.28}
\end{equation*}
$$

which shows the relative strength of convection and diffusion and must not exceed certain value, specific for each scheme.

The table with discretization schemes widely used in CFD with brief explanation of their advantages and disadvantages is shown below in "Table 8". (Versteeg \& Malalasekera 2007, 143-178).

Table 8. Discretization schemes, commonly used in CFD.

| Discretization <br> scheme | advantages | disadvantages |
| :---: | :---: | :---: |
| Central <br> differencing | The simplest formulation of <br> coefficients "a". | Works only for <br> $P e<2$. |
| Upwind <br> differencing | No limitations on Peclet number. | Only first order Taylor <br> series truncation <br> error. Problem of <br> "false diffusion". |
| Hybrid | Switches from upwind to central <br> differencing scheme, when <br> Pe $<2$, hence more suitable for <br> modelling of boundary layer. | Same as for upwind <br> differencing scheme <br> when Pe $\geq 2$. |
| Power-law | Accuracy is higher than in Hybrid <br> scheme. | Only applicable to <br> one-dimensional <br> flows. |
| QUICK <br> (Quadratic <br> upstream <br> interpolation <br> for | Accuracy is higher than in Hybrid <br> scheme. Resultant false diffusion <br> is smaller than in hybrid scheme <br> too. | Minor under- and <br> overshoots in <br> solution, that can give <br> physically unrealistic <br> results (e.g. negative <br> kinetics) |
| TVD <br> (Total <br> variation <br> diminishing) | Complete elimination of under- <br> and overshoots. Same accuracy <br> as in QUICK scheme, resulting in <br> the most accurate and realistic <br> results. | Mathematical <br> complexity |

### 2.4.3 Pressure-velocity coupling

As it was mentioned before, momentum equations contain non-linear convective terms (e.g. unknown velocity component multiplied by its spatial derivative), what causes need to use iterative techniques (Like Newton-Raphson method) to obtain solution with sufficient accuracy. However, nonlinear terms are not the only reason to use iterative methods to solve governing equations. Unlike compressible flows, where pressure can be obtained from solving corresponding transport equation of energy (temperature) and certain gas law, incompressible flows don't
have such relation with density and temperature, but application of correct pressure field function must yield to satisfied continuity. Therefore, one can find correct pressure function iteratively from initially guessed function, by performing certain amount of iterations, correcting the "guessed" pressure function, until continuity equations turns out to be sufficiently satisfied (converge). (Versteeg \& Malalasekera 2007, 179 196).

In order to couple the pressure and velocity to convection-diffusion equation, one has to discretize the pressure, first. There exist several methods to accomplish this problem, however only one way, called staggered grid arrangement, allows to obtain sufficiently realistic results. The pressure gradients from 2.2.25 in staggered arrangement are defined as follows:

$$
\begin{gather*}
\frac{\partial p}{\partial x}=\frac{p_{P}-p_{W}}{\delta x_{W P}} \\
\frac{\partial p}{\partial y}=\frac{p_{P}-p_{S}}{\delta y_{S P}} \\
\frac{\partial p}{\partial z}=\frac{p_{P}-p_{B}}{\delta z_{B P}} \tag{2.4.29}
\end{gather*}
$$

Recalling the formula 2.4.13 that can also be applied to convectiondiffusion problems (coefficients " a " will vary depending on discretization scheme), applied to $x$-momentum equation with subtraction of corresponding pressure gradient will result to discretized equation of $x$ momentum:

$$
\begin{equation*}
a_{P} u_{P}=\sum a_{n b} u_{n b}-\frac{p_{P}-p_{W}}{\delta x_{W P}} \Delta V_{P}+\bar{S} \Delta V_{P} \tag{2.4.30}
\end{equation*}
$$

If one assumes uniform square grid (two-dimensional problem), index $P$ corresponding to target node (cell centre) can be represented by double index " J ", where " l " corresponds to certain row of mesh and " J " - to column. In the same way, one also has to use lower case letters to refer to points at cell faces, where " $i$ " corresponds to certain row, and " $j$ " - to column (See the Figure 12). With this notation 2.4 .31 will change to:

$$
a_{i J} u_{i J}=\sum a_{n b} u_{n b}-\frac{p_{I J}-p_{(I-1) J}}{\delta x_{W P}} \Delta V_{I J}+\bar{S} \Delta V_{I J}
$$

And further to:

$$
\begin{equation*}
a_{i j} u_{i J}=\sum a_{n b} u_{n b}+\left(p_{(I-1) J}-p_{I J}\right) A_{i J}+b_{i J} \tag{2.4.31}
\end{equation*}
$$

The physical meaning of such model is that all scalar quantities (density, pressure, temperature, etc.) are computed exactly at nodal points, while velocity components are evaluated at cell faces, what allows properly catch non-uniform behaviour of field functions if it is the case. (Versteeg \& Malalasekera 2007, 180-186)
As previously mentioned, in order to solve pressure and velocity functions, one has to specify initially guessed value, that will be later corrected to more realistic result using iterative methods, so the first step in solving flow equations is the specification of guessed value, which can be obtained from decomposition of correct values:

$$
\begin{align*}
& p=p^{*}+p^{\prime} \\
& u=u^{*}+u^{\prime} \\
& v=v^{*}+v^{\prime} \\
& w=w^{*}+w^{\prime} \tag{2.4.32}
\end{align*}
$$

Where upper script "*" denotes the initially stated (guessed) value, and ' denotes the value correction, necessary to satisfy governing equations. (Versteeg \& Malalasekera 2007, 186-190).


Figure 16. "IJ" and "ij" notation of uniform two-dimensional grid with illustrated application of continuity equation (Versteeg \& Malalasekera 2007, 188).

There exist four algorithms widely used in CFD applications:

- SIMPLE (Semi-Implicit Method for Pressure-Linked Equations)
- SIMPLER (SIMPLE Revised)
- SIMPLEC (SIMPLE Consistent)
- PISO (Pressure Implicit with Splitting of Operations)

In addition to step sequence, they vary by equations used to determine the corrections " $\phi$ '". The actual schemes of SIMPLE, SIMPLEC and PISO are shown in Figure 17, 18 and 19.


Figure 17. SIMPLE algorithm diagram (Versteeg \& Malalasekera 2007, 190).


Figure 18. SIMPLER algorithm diagram (Versteeg \& Malalasekera 2007, 192).


Figure 19. PISO algorithm diagram (Versteeg \& Malalasekera 2007, 195).

Where upper script "**" denotes corrected pressure, "'"" denotes second correction and "***" denotes twice-corrected pressure:

$$
\begin{equation*}
p^{* * *}=p^{* *}+p^{\prime \prime}=p^{*}+p^{\prime}+p^{\prime \prime} \tag{2.4.33}
\end{equation*}
$$

And for " $d$ " with corresponding indices:

$$
\begin{equation*}
d=\frac{A}{a} \tag{2.4.34}
\end{equation*}
$$

Every listed algorithm has own advantages and disadvantages. Particularly SIMPLE algorithm executes the least amount of computations per iteration, but requires more iterations to converge, while SIMPLER algorithm shows opposite behaviour due to better pressure correction scheme. On the other hand PISO shows better performance when there is no coupling of momentum equations with scalar variables (incompressible flows modelled using LES or DNS, laminar flows), while SIMPLER and SIMPLEC show better performance with highly coupled flows (compressible flows with RANS models). (Versteeg \& Malalasekera 2007, 196).

## 3 EXAMPLE SOLVED IN MATHCAD 15

As one can notice, mathematical methods used in CFD are extremely lengthy to solve. The aim of examples this chapter is to show the application of theory, covered in Chapter 2 and demonstrate the need in such mathematical software packages as ANSYS or COMSOL: complexity of implementation of covered mathematical methods for extremely simple one-dimensional pressure-driven flow with coarse mesh.

### 3.1 Problem specification

As an example, one can consider steady, inviscid, incompressible pressure-driven flow in converging-diverging nozzle, shown in figure 20.


Figure 20. Converging-diverging nozzle with marked nodal points for pressure field (Arabic numerals) and grid for velocity field (Roman numerals).

Note that cross-sectional area is assumed to change linearly along symmetry axis until point 4 , where it starts to stay constant. The applied solution scheme will be upwind-differencing scheme, and used algorithm - SIMPLE.

Problem specifications:
Area at inlet: $0.5 \mathrm{~m}^{2}$
Area at outlet: $0.4 \mathrm{~m}^{2}$
Area at point 3: $0.3 \mathrm{~m}^{2}$
Outlet pressure: atmospheric pressure $p_{a}=101325$ ( Pa )
Inlet pressure: $p_{\text {in }}=2 p_{a}$
Overall length: 1.6 m
Fluid density: $1000 \mathrm{~kg} / \mathrm{m}^{3}$
Since the flow is simplified to incompressible, steady, one-dimensional and inviscid, one can use a Bernoulli's equation to find fluid velocity in the outlet and corresponding simplified continuity equation to find velocity and pressures at all nodal points:

$$
\begin{equation*}
p+\rho \frac{u^{2}}{2}=\text { const } \tag{3.1.1}
\end{equation*}
$$

$$
\begin{equation*}
u A=\text { const } \tag{3.1.2}
\end{equation*}
$$

With the origin placed at the point 1, one can get the following results:

Table 9. Exact solution, using Bernoulli's equation.

| Corresponding <br> nodal point | $\mathrm{x}(\mathrm{m})$ | $\mathrm{A}\left(\mathrm{m}^{2}\right)$ | $\mathrm{p}(\mathrm{Pa})$ | $\mathrm{u}(\mathrm{m} / \mathrm{s})$ |
| :---: | :---: | :---: | :---: | :---: |
| 1 | 0 | 0.5 | 137802 | 11.388415 |
| I | 0.2 | 0.45 | 122590.740741 | 12.653795 |
| 2 | 0.4 | 0.4 | 101325 | 14.235519 |
| II | 0.6 | 0.35 | 70307.142857 | 16.269165 |
| 3 | 0.8 | 0.3 | 22516.666667 | 18.980692 |
| III | 1 | 0.35 | 70307.142857 | 16.269165 |
| 4 | 1.2 | 0.4 | 101325 | 14.235519 |
| IV | 1.4 | 0.4 | 101325 | 14.235519 |
| 5 | 1.6 | 0.4 | 101325 | 14.235519 |

These results will be used later for comparison with numerical method.

### 3.2 Numerical solution

1) According to SIMPLE algorithm one has firstly specify initial pressure and velocity fields.

For pressure field nodal points one cane make the next guess:
Table 10. Initial guess for pressure field.

| 1 | 2 | 3 | 4 | 5 |
| :---: | :---: | :---: | :---: | :---: |
| $1.5 p_{a}$ | $p_{a}$ | 0 | $p_{a}$ | $p_{a}$ |

For velocity field one can make a guess of volumetric flow at point IV and use continuity equation to find the rest "guessed" values. Say let the velocity at point IV to be $14(\mathrm{~m} / \mathrm{s})$, then one can obtain the following table:

Table 11. Initial guess for velocity field.

| I | II | III | IV |
| :---: | :---: | :---: | :---: |
| $12.444444(\mathrm{~m} / \mathrm{s})$ | $16(\mathrm{~m} / \mathrm{s})$ | $16(\mathrm{~m} / \mathrm{s})$ | $14(\mathrm{~m} / \mathrm{s})$ |

2) With given simplifications governing equations (continuity and momentum conservation correspondingly) will transform to:

$$
\begin{equation*}
\frac{d}{d x}(\rho A u)=0 \tag{3.2.1}
\end{equation*}
$$

$$
\begin{equation*}
\rho u A \frac{d u}{d x}=-A \frac{d p}{d x} \tag{3.2.2}
\end{equation*}
$$

After discretization of second equation one obtains the next form of momentum equation:

$$
\begin{equation*}
(\rho u A)_{e} u_{e}-(\rho u A)_{w} u_{w}=\frac{p_{w}-p_{e}}{\Delta x} \Delta V \tag{3.2.3}
\end{equation*}
$$

For a single nodal point " $P$ " the momentum equation will have $a$ following form:

$$
\begin{equation*}
a_{P} u_{P}^{*}=a_{W} u_{W}^{*}+a_{E} u_{E}^{*}+S_{u} \tag{3.2.4}
\end{equation*}
$$

The coefficients " a " are obtained from applied solution scheme, which is the upwind difference scheme in this case. For this solution scheme coefficients " a " are listed in tables below

Table 12. Coefficients "a" for upwind differencing solution scheme.

| $a_{P}$ | $a_{W}+a_{E}+a_{S}+a_{N}+a_{B}+a_{T}+\Delta F$ |
| :---: | :---: |
| $a_{W}$ | $D_{w}+\max \left(F_{w}, 0\right)$ |
| $a_{E}$ | $D_{e}+\max \left(-F_{e}, 0\right)$ |
| $a_{S}$ | $D_{S}+\max \left(F_{s}, 0\right)$ |
| $a_{N}$ | $D_{n}+\max \left(-F_{n}, 0\right)$ |
| $a_{B}$ | $D_{b}+\max \left(F_{b}, 0\right)$ |
| $a_{T}$ | $D_{t}+\max \left(-F_{t}, 0\right)$ |
| $\Delta F$ | $F_{e}-F_{w}+F_{n}-F_{s}+F_{t}-F_{b}$ |

Table 13. Values of " $F$ " and " $D$ ".

| Face | w | e | s | n | b | t |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| F | $(\rho u)_{w} A_{w}$ | $(\rho u)_{e} A_{e}$ | $(\rho v)_{s} A_{S}$ | $(\rho v)_{n} A_{n}$ | $(\rho w)_{b} A_{b}$ | $(\rho w)_{t} A_{t}$ |
| D | $\frac{\Gamma_{w}}{\delta x_{W P}} A_{w}$ | $\frac{\Gamma_{e}}{\delta x_{P E}} A_{e}$ | $\frac{\Gamma_{s}}{\delta y_{S P}} A_{s}$ | $\frac{\Gamma_{n}}{\delta y_{P N}} A_{n}$ | $\frac{\Gamma_{b}}{\delta z_{B P}} A_{b}$ | $\|$$\frac{\Gamma_{t}}{\delta z_{P T}} A_{t}$ |

Since the flow was assumed to be one-dimensional all coefficients that contribute to south, north, top and bottom coefficients "a" will be equal to zero. Moreover, the flow was assumed to be inviscid (viscosity equals zero), therefore all values of diffusive terms " $D$ " will be equal zero too.

Source term $S_{u}$ in 3.2.4 hides the pressure gradient term of discretized momentum equation 3.2.3:

$$
\begin{equation*}
S_{u}=\frac{p_{w}-p_{e}}{\Delta x} \Delta V=\frac{1}{2}\left(p_{w}-p_{e}\right)\left(A_{w}+A_{e}\right)=\left(p_{w}-p_{e}\right) A_{P} \tag{3.2.5}
\end{equation*}
$$

Finally, after all simplifications the coefficients for 3.2.4 are:

$$
\begin{gathered}
a_{P}=a_{W}+a_{E}+\left(F_{e}-F_{w}\right) \\
F_{w}=\rho u_{w} A_{w} \\
F_{e}=\rho u_{e} A_{e} \\
a_{W}=F_{w} \\
a_{E}=0
\end{gathered}
$$

$$
\begin{equation*}
S_{u}=\frac{1}{2}\left(p_{w}-p_{e}\right)\left(A_{w}+A_{e}\right)=\left(p_{w}-p_{e}\right) A_{P} \tag{3.2.6}
\end{equation*}
$$

The parameter " d " necessary for pressure correction equation is determined as follows:

$$
\begin{equation*}
d_{P}=\frac{A_{P}}{a_{P}} \tag{3.2.7}
\end{equation*}
$$

3) The discretized form of 3.2 .1 will be:

$$
\begin{gather*}
(\rho u A)_{e}-(\rho u A)_{w}=0  \tag{3.2.8}\\
a_{P} p_{P}^{\prime}=a_{W} p_{W}^{\prime}+a_{E} p_{E}^{\prime}+b^{\prime} \tag{3.2.9}
\end{gather*}
$$

Where:

$$
\begin{gather*}
a_{P}=a_{W}+a_{E} \\
a_{W}=(\rho d A)_{w} \\
a_{E}=(\rho d A)_{e} \\
b^{\prime}=F_{w}^{*}-F_{e}^{*}=\left(\rho u^{*} A\right)_{w}-\left(\rho u^{*} A\right)_{e} \tag{3.2.10}
\end{gather*}
$$

The key approximation that is the core of SIPMPLE algorithm is the following formula for velocity correction that is obtained from pressure corrections:

$$
\begin{equation*}
u^{\prime}=d\left(p_{I}^{\prime}-p_{I+1}^{\prime}\right) \tag{3.2.11}
\end{equation*}
$$

Once pressure and velocity corrections are found one can find the solution of pressure and velocity fields, by simple summation of guessed values with corrections as shown in 2.4.32.
4) The next step is to solve velocity field and correction parameters " $d$ " using the guessed values. To accomplish this, one has to form the system of linear equations, using 3.2.4, 3.2.6 and 3.2.7.

The solution in Mathcad 15 begins with specification of coefficients "a" and source terms of momentum equations for velocity nodes II and III as follows:

$$
\begin{aligned}
& \mathrm{p}_{\mathrm{a}}:=10132: \quad \mathrm{u}_{\mathrm{I}}:=12.44444 \quad \mathrm{u}_{\mathrm{II}}:=1 \mathrm{t} \quad \mathrm{u}_{\text {III }}:=1 \mathrm{t} \quad \mathrm{u}_{\text {IV }}:=14 \\
& \mathrm{~A}_{\mathrm{I}}:=0.45 \quad \mathrm{~A}_{\mathrm{II}}:=0.35 \quad \mathrm{~A}_{\mathrm{III}}:=0.35 \quad \mathrm{~A}_{\mathrm{IV}}:=0.4 \quad \rho:=100( \\
& \mathrm{A}_{1}:=0.5 \quad \mathrm{~A}_{2}:=0.4 \quad \mathrm{~A}_{3}:=0.3 \quad \mathrm{~A}_{4}:=0.4 \quad \mathrm{~A}_{5}:=0.4 \\
& \mathrm{p}_{1}:=1.5 \cdot \mathrm{p}_{\mathrm{a}} \quad \mathrm{p}_{2}:=\mathrm{p}_{\mathrm{a}} \quad \mathrm{p}_{3}:=0 \quad \mathrm{p}_{4}:=\mathrm{p}_{\mathrm{a}} \quad \mathrm{p}_{5}:=\mathrm{p}_{\mathrm{a}} \\
& \mathrm{~F}_{\mathrm{wII}}:=\rho \cdot \mathrm{A}_{2} \cdot \frac{1}{2} \cdot\left(\mathrm{u}_{\mathrm{II}}+\mathrm{u}_{\mathrm{I}}\right) \quad \mathrm{F}_{\mathrm{eII}}:=\rho \cdot \mathrm{A}_{3} \cdot \frac{1}{2} \cdot\left(\mathrm{u}_{\mathrm{II}}+\mathrm{u}_{\mathrm{III}}\right) \\
& { }^{a_{W I I}}:=\mathrm{F}_{\text {wII }} \quad{ }^{a_{E I I}}:=0 \\
& \mathrm{a}_{\text {II }}:=\mathrm{a}_{\mathrm{WII}}+\mathrm{a}_{\text {EII }}+\left(\mathrm{F}_{\text {eII }}-\mathrm{F}_{\mathrm{wII}}\right) \\
& \mathrm{S}_{\text {II }}:=\left(\mathrm{p}_{2}-\mathrm{p}_{3}\right) \cdot \mathrm{A}_{\text {II }} \\
& \mathrm{d}_{\text {II }}:=\frac{\mathrm{A}_{\text {II }}}{\mathrm{a}_{\text {II }}} \\
& \mathrm{F}_{\text {wIII }}:=\rho \cdot \mathrm{A}_{3} \cdot \frac{1}{2} \cdot\left(\mathrm{u}_{\text {III }}+\mathrm{u}_{\text {II }}\right) \quad \mathrm{F}_{\text {eIII }}:=\rho \cdot \mathrm{A}_{4} \cdot \frac{1}{2} \cdot\left(\mathrm{u}_{\text {III }}+\mathrm{u}_{\mathrm{IV}}\right) \\
& { }^{a_{W I I I}}:=\mathrm{F}_{\text {wIII }} \quad{ }^{a_{E I I I}}:=0 \\
& \mathrm{a}_{\text {III }}:=\mathrm{a}_{\text {WIII }}+\mathrm{a}_{\text {EIII }}+\left(\mathrm{F}_{\text {eIII }}-\mathrm{F}_{\text {wIII }}\right) \\
& \mathrm{S}_{\text {III }}:=\left(\mathrm{p}_{3}-\mathrm{p}_{4}\right) \cdot \mathrm{A}_{\text {III }} \\
& \mathrm{d}_{\text {III }}:=\frac{\mathrm{A}_{\text {III }}}{\mathrm{a}_{\text {III }}}
\end{aligned}
$$

The next step is specifying the same coefficients for momentum equations of nodes I and IV. They are placed on boundaries of control volume of problem and require the special treatment.

Due to absence of the neighbouring node for the velocity node " $I$ ", one has to make the correlation with the pressure node " 1 " by specifying the velocity at pressure node " 1 ", what can be done using the Bernoulli's (3.1) and continuity (3.2) equation as follows:

$$
\begin{equation*}
p_{\text {in }}=p_{1}+\frac{1}{2} \rho u_{1}^{2} \tag{3.2.12}
\end{equation*}
$$

$$
\begin{equation*}
u_{I} A_{I}=u_{1} A_{1} \tag{3.2.13}
\end{equation*}
$$

The combination of these two equations will yield to:

$$
\begin{equation*}
p_{1}=p_{\text {in }}-\frac{1}{2} \rho u_{I}^{2}\left(\frac{A_{I}}{A_{1}}\right)^{2} \tag{3.2.14}
\end{equation*}
$$

The discretized momentum equation for velocity node " 1 " is determined as follows:

$$
\begin{equation*}
F_{e} u_{I}-F_{w} u_{1}=\left(p_{1}-p_{2}\right) A_{I} \tag{3.2.15}
\end{equation*}
$$

Where $F_{w}=\rho u_{1} A_{1}=\rho u_{I} A_{I}$
The substitution of 3.15 and 3.16 to 3.17 will result to:

$$
\begin{equation*}
\left(F_{e}-F_{w}\left(\frac{A_{I}}{A_{1}}\right)+\frac{1}{2} F_{w}\left(\frac{A_{I}}{A_{1}}\right)^{2}\right) u_{1}=\left(p_{i n}-p_{2}\right) A_{I} \tag{3.2.16}
\end{equation*}
$$

In order to avoid negative value for coefficient " $a_{I}$ " one has to rearrange 3.18 to next form:

$$
\begin{equation*}
\left(F_{e}+\frac{1}{2} F_{w}\left(\frac{A_{I}}{A_{1}}\right)^{2}\right) u_{1}=\left(p_{\text {in }}-p_{2}\right) A_{I}+F_{w}\left(\frac{A_{I}}{A_{1}}\right) u_{I}^{*} \tag{3.2.17}
\end{equation*}
$$

Where $u_{I}^{*}$ denotes velocity at point "I" from previous iteration. Therefore " $a_{I}$ ", " $S_{I}$ " and " $d_{I}$ " are determined as follows:

$$
\begin{aligned}
& \mathrm{p}_{\mathrm{in}}:=2 \cdot \mathrm{p}_{\mathrm{a}} \quad \mathrm{~F}_{\mathrm{wI}}:=\rho \cdot \mathrm{A}_{\mathrm{I}} \cdot \mathrm{u}_{\mathrm{I}} \quad \mathrm{~F}_{\mathrm{eI}}:=\rho \cdot \mathrm{A}_{2} \cdot \frac{1}{2} \cdot\left(\mathrm{u}_{\mathrm{I}}+\mathrm{u}_{\mathrm{II}}\right) \\
& \mathrm{a}_{\mathrm{I}}:=\mathrm{F}_{\mathrm{eI}}+\mathrm{F}_{\mathrm{wI}} \cdot \frac{1}{2} \cdot\left(\frac{\mathrm{~A}_{\mathrm{I}}}{\mathrm{~A}_{1}}\right)^{2} \\
& \mathrm{~S}_{\mathrm{I}}:=\left(\mathrm{p}_{\mathrm{in}}-\mathrm{p}_{2}\right) \cdot \mathrm{A}_{\mathrm{I}}+\mathrm{F}_{\mathrm{wI}} \cdot\left(\frac{\mathrm{~A}_{\mathrm{I}}}{\mathrm{~A}_{1}}\right)^{2} \cdot \mathrm{u}_{\mathrm{I}} \quad \quad \mathrm{~d}_{\mathrm{I}}:=\frac{\mathrm{A}_{\mathrm{I}}}{\mathrm{a}_{\mathrm{I}}}
\end{aligned}
$$

For the velocity node "IV":

$$
\begin{aligned}
\mathrm{F}_{\mathrm{wIV}} & :=\rho \cdot\left[\frac{\left(\mathrm{u}_{\mathrm{III}}+\mathrm{u}_{\mathrm{IV}}\right)}{2}\right] \cdot \mathrm{A}_{4} \quad \mathrm{~F}_{\mathrm{eIV}}:=\rho \cdot \mathrm{u}_{\mathrm{IV}} \mathrm{~A}_{\mathrm{IV}} \\
\mathrm{a}_{\mathrm{WIV}} & :=\mathrm{F}_{\mathrm{wIV}} \quad \mathrm{a}_{\mathrm{EIV}}:=0 \\
\mathrm{a}_{\mathrm{IV}} & :=\mathrm{a}_{\mathrm{WIV}}+\mathrm{a}_{\mathrm{EIV}}+\left(\mathrm{F}_{\mathrm{eIV}}-\mathrm{F}_{\mathrm{wIV}}\right) \\
\mathrm{S}_{\mathrm{IV}} & :=\left(\mathrm{p}_{4}-\mathrm{p}_{5}\right) \cdot \mathrm{A}_{\mathrm{IV}} \\
\mathrm{~d}_{\mathrm{IV}} & :=\frac{\mathrm{A}_{\mathrm{IV}}}{\mathrm{a}_{\mathrm{IV}}}
\end{aligned}
$$

Once coefficients for all four momentum (velocity) nodes are determined, one is ready to solve the system of four algebraic equations with four unknowns, where unknowns are computed velocities at momentum nodes denoted with capital " U " s :

Given

$$
\left(\begin{array}{c}
\mathrm{U}_{\mathrm{I}} \\
\mathrm{U}_{\text {II }} \\
\mathrm{U}_{\text {III }} \\
\mathrm{U}_{\text {IV }}
\end{array}\right):=\operatorname{Find}\left(\mathrm{U}_{\mathrm{I}}, \mathrm{U}_{\text {II }}, \mathrm{U}_{\text {III }}, \mathrm{U}_{\text {IV }}\right) \rightarrow\left(\begin{array}{l}
12.82464158689713503 \\
22.58785622660694547 \\
12.1596599812855563 \\
13.0282071228059532 \phi
\end{array}\right.
$$

These computed velocities will be used to compute the pressure corrections.
5) For pressure nodes " 2 ", " 3 " and " 4 " pressure correction equations are determined directly from 3.11:

For nodes " 1 " and " 5 " the pressure corrections are set to 0 . Therefore one gets the following system of equations for pressure corrections:

$$
\begin{aligned}
& \mathrm{p}_{1}^{\prime}:=0 \quad \mathrm{p}_{5}^{\prime}:=0 \\
& \text { Given } \\
& \mathrm{a}_{2} \cdot \mathrm{p}_{2}^{\prime}=\mathrm{a}_{2} \mathrm{~W} \cdot \mathrm{p}^{\prime}{ }_{1}+\mathrm{a}_{2 \mathrm{E}} \mathrm{p}_{3}^{\prime}+\mathrm{b}_{2}^{\prime} \\
& \mathrm{a}_{3} \cdot \mathrm{p}_{3}^{\prime}=\mathrm{a}_{3}{ }^{W} \cdot \mathrm{p}_{2}^{\prime}+\mathrm{a}_{3 \mathrm{E}^{2}} \mathrm{p}_{4}^{\prime}+\mathrm{b}_{3}^{\prime} \\
& \mathrm{a}_{4} \cdot \mathrm{p}_{4}^{\prime}=\mathrm{a}_{4} \mathrm{~W}^{\prime} \cdot \mathrm{p}_{3}^{\prime}+\mathrm{a}_{4 \mathrm{E}^{\mathrm{E}} \mathrm{p}_{5}^{\prime}+\mathrm{b}_{4}^{\prime}}
\end{aligned}
$$

$$
\begin{aligned}
& \mathrm{a}_{2 \mathrm{~W}}:=\rho \cdot \mathrm{d}_{\mathrm{I}} \cdot \mathrm{~A}_{\mathrm{I}} \quad \mathrm{a}_{2 \mathrm{E}}:=\rho \cdot \mathrm{d}_{\text {II }} \cdot \mathrm{A}_{I I} \\
& F_{2 w}:=\rho \cdot U_{I} \cdot A_{I} \quad F_{2 E}:=\rho \cdot U_{I I} \cdot A_{I I} \\
& \mathrm{a}_{2}:=\mathrm{a}_{2 \mathrm{~W}}+\mathrm{a}_{2 \mathrm{E}} \quad \mathrm{~b}_{2}:=\mathrm{F}_{2 \mathrm{w}}-\mathrm{F}_{2 \mathrm{E}} \\
& a_{3 W}:=\rho \cdot d_{I I} A_{I I} \quad a_{3 E}:=\rho \cdot d_{I I I}{ }^{-A_{I I I}} \\
& \mathrm{~F}_{3 \mathrm{w}}:=\rho \cdot \mathrm{U}_{\text {II }} \cdot \mathrm{A}_{\text {II }} \quad \mathrm{F}_{3 \mathrm{E}}:=\rho \cdot \mathrm{U}_{\text {III }} \cdot \mathrm{A}_{\text {III }} \\
& a_{3}:=a_{3 W}+a_{3 E} \quad b_{3}^{\prime}:=F_{3 w}-F_{3 E} \\
& \mathrm{a}_{4 \mathrm{~W}}:=\rho \cdot \mathrm{d}_{\text {III }} \cdot \mathrm{A}_{\text {III }} \quad \mathrm{a}_{4 \mathrm{E}}:=\rho \cdot \mathrm{d}_{\text {IV }} \mathrm{A}_{\text {IV }} \\
& \mathrm{F}_{4 \mathrm{w}}:=\rho \cdot \mathrm{U}_{\text {III }} \cdot \mathrm{A}_{\text {III }} \quad \mathrm{F}_{4 \mathrm{E}}:=\rho \cdot \mathrm{U}_{\text {IV }} \mathrm{A}_{\text {IV }} \\
& \mathrm{a}_{4}:=\mathrm{a}_{4 \mathrm{~W}}+\mathrm{a}_{4 \mathrm{E}} \quad \mathrm{~b}_{4}:=\mathrm{F}_{4 \mathrm{w}}-\mathrm{F}_{4 \mathrm{E}}
\end{aligned}
$$

$$
\begin{aligned}
& { }^{\mathrm{a}} \mathrm{I}_{\mathrm{I}} \cdot \mathrm{U}_{\mathrm{I}}=\mathrm{S}_{\mathrm{I}} \\
& { }^{a_{I I}} \mathrm{U}_{\text {II }}={ }^{\mathrm{a}_{\text {WII }}} \mathrm{U}_{\mathrm{I}}+\mathrm{S}_{\text {II }} \\
& \mathrm{a}_{\text {III }} \cdot \mathrm{U}_{\text {III }}=\mathrm{a}_{\text {WIII }} \cdot \mathrm{U}_{\text {II }}+\mathrm{S}_{\text {III }} \\
& \mathrm{a}_{\text {IV }} \mathrm{U}_{\text {IV }}=\mathrm{a}_{\text {WIV }} \cdot \mathrm{U}_{\text {III }}+\mathrm{S}_{\text {IV }}
\end{aligned}
$$

$$
\left(\begin{array}{l}
\mathrm{p}_{2}^{\prime} \\
\mathrm{p}_{3}^{\prime} \\
\mathrm{p}_{4}^{\prime}
\end{array}\right):=\operatorname{Find}\left(\mathrm{p}_{2}^{\prime}, \mathrm{p}_{3}^{\prime}, \mathrm{p}_{4}^{\prime}\right) \rightarrow\left(\begin{array}{l}
2458.265191215116959 \\
88553.5315582056770 才 \\
17403.5360257200641
\end{array}\right)
$$

Finally, computed velocities and pressures are:

$$
\begin{aligned}
& \mathrm{u}_{\text {Icomp }}:=\mathrm{U}_{\mathrm{I}}+\mathrm{d}_{\mathrm{I}} \cdot\left(\mathrm{p}_{1}-\mathrm{p}_{2}^{\prime}\right)=12.68561496793672 \\
& \mathrm{u}_{\text {IIcomp }}:=\mathrm{U}_{\mathrm{II}}+\mathrm{d}_{\text {II }}\left(\mathrm{p}_{2}^{\prime}-\mathrm{p}_{3}\right)=16.31007638734721 \epsilon \\
& \mathrm{u}_{\text {IIIcomp }}:=\mathrm{U}_{\text {IIII }}+\mathrm{d}_{\text {IIII }} \cdot\left(\mathrm{p}_{3}{ }_{3}-\mathrm{p}_{4}{ }_{4}\right)=16.31007638734721 \epsilon \\
& \mathrm{u}_{\text {IVcomp }}:=\mathrm{U}_{\mathrm{IV}}+\mathrm{d}_{\mathrm{IV}}\left(\mathrm{p}_{4}^{\prime}-\mathrm{p}_{5}^{\prime}\right)=14.27131683892881 \text { ร } \\
& \mathrm{p}_{1 \text { comp }}:=\mathrm{p}_{1}+\mathrm{p}_{1}=151987.5 \\
& \mathrm{p}_{2 \mathrm{comp}}:=\mathrm{p}_{2}+\mathrm{p}_{2}^{\prime}=103783.2651912151^{\prime} \\
& \left.\mathrm{p}_{3 \text { comp }}:=\mathrm{p}_{3}+\mathrm{p}_{3}^{\prime}=88553.5315582056\right\} \\
& \mathrm{p}_{4 \text { comp }}:=\mathrm{p}_{4}+\mathrm{p}_{4}^{\prime}=118728.5360257200^{-} \\
& \mathrm{p}_{5 \mathrm{comp}}:=\mathrm{p}_{5}+\mathrm{p}_{5}^{\prime}=101325
\end{aligned}
$$

The last step in SIMPLE algorithm is the usage of so-called relaxation, which is defined as:

$$
u_{\text {new }}=(1-\alpha) u_{\text {old }}+\alpha u_{\text {comp }}
$$

$$
\begin{equation*}
p_{\text {new }}=(1-\alpha) p_{\text {old }}+\alpha p_{\text {comp }} \tag{3.2.18}
\end{equation*}
$$

Where " $\alpha$ " is so-called relaxation coefficient defined by programme or user and always less than one.

Particularly for $\alpha=0.75$ :

$$
\begin{aligned}
& \alpha:=0.7^{〔} \\
& \mathrm{u}_{\text {Inew }}:=(1-\alpha) \cdot \mathrm{u}_{\mathrm{I}}+\alpha \cdot \mathrm{u}_{\text {Icomp }}=12.62532222595254 \\
& \mathrm{u}_{\text {IInew }}:=(1-\alpha) \cdot \mathrm{u}_{\text {II }}+\alpha \cdot \mathrm{u}_{\text {IIcomp }}=16.23255729051041 \\
& \mathrm{u}_{\text {IIInew }}:=(1-\alpha) \cdot \mathrm{u}_{\text {III }}+\alpha \cdot \mathrm{u}_{\text {IIIcomp }}=16.23255729051041 \\
& \mathrm{u}_{\text {IVnew }}:=(1-\alpha) \cdot \mathrm{u}_{\text {IV }}+\alpha \cdot \mathrm{u}_{\text {IVcomp }}=14.20348762919661 \\
& \mathrm{p}_{1 \text { new }}:=(1-\alpha) \cdot \mathrm{p}_{1}+\alpha \cdot \mathrm{p}_{1 \text { comp }}=151987.5 \\
& \mathrm{p}_{2 \text { new }}:=(1-\alpha) \cdot \mathrm{p}_{2}+\alpha \cdot \mathrm{p}_{2 \text { comp }}=103168.6988934113 ؛ \\
& \mathrm{p}_{3 \text { new }}:=(1-\alpha) \cdot \mathrm{p}_{3}+\alpha \cdot \mathrm{p}_{3 \text { comp }}=66415.1486686542 t \\
& \mathrm{p}_{\text {4new }}:=(1-\alpha) \cdot \mathrm{p}_{4}+\alpha \cdot \mathrm{p}_{4 \text { comp }}=114377.6520192900 ؛ \\
& \mathrm{p}_{5 \text { new }}:=(1-\alpha) \cdot \mathrm{p}_{5}+\alpha \cdot \mathrm{p}_{5 \text { comp }}=101325
\end{aligned}
$$

The last set of values is the outcome of a single iteration in SIMPLE algorithm. In order to execute the second iteration one simply has to set the final result of first iteration as guess values for second iteration and repeat the procedure. Results (except pressures at nodes " 1 " (because in most CFD codes this value will be reserved for stagnation pressure at pressure inlet, which is double atmospheric pressure in this case must not change with any number of performed iterations) and " 3 ") show some agreement with the exact solution in Table 9 even though they not necessary should, since the grid consist only from five pressure a four momentum (velocity) nodes, that makes it to be a relatively coarse mesh. The accuracy can be improved by the grid refinement (increasing the amount of nodes). But one must remember that this will increase the complexity of mathematical model and require more iterations to converge (Versteeg \& Malalasekera 2007, 209-210).

The disadvantage of poor accuracy in numerical methods is compensated by their ability to be expanded to more complex three-dimensional, compressible, viscous and even ionized flows, what together with the absence of alternatives ensured the popularity of CFD in industrial and scientific applications.

Considered example explicitly shows how lengthy the computation of flow can be even with coarse mesh. Therefore, in order to handle more complex scenarios such as three-dimensional transient flow, it's strongly suggested to use either commercial software like ANSYS or COMSOL or any free CFD software, that can be found here:
https://www.cfd-online.com/Wiki/Codes.
In order to learn how to use ANSYS FLUENT one can go to link below which contains tutorials in FLUENT covering all aspects, necessary to know for modelling of complex flows.
Link to ANSYS FLUENT tutorials:
https://confluence.cornell.edu/display/SIMULATION/FLUENT+Learning+ Modules

Once tutorials are completed, one can continue the self-study by exploring the following book, which also gives the insight into programming aspects of CFD: Moukalled F., Mangani L., Darwish M. (2016) The Finite Volume Method in Computational Fluid Dynamics: An Advanced Introduction with OpenFOAM and Matlab. Stuttgart: Springer.

## 4 PI-THEOREM HANDOUTS

CFD provides Mechanical Engineering designers with an inexpensive alternative to actual experiments supported by a high flexibility in changing of prototype design. However, once the results from several simulations are obtained, one has to build up satisfactory mathematical models. The main difficulty at this stage is usually a wide range of variables involved in the simulated process including: velocity at inlet, the density and viscosity of the fluid, ambient temperature, thermal conductivity, drag force, geometry of control volume, frequency of vibration, enthalpy etc., making the question of which variables are dependent on each other to be difficult to handle. To answer this question completely, one can use the Bucking-ham Pi-Theorem (or simply Pi-Theorem). The Pi-Theorem states that any set of variables can be changed to the set of distinct non-dimensional groups, that are dependent on each other. The exact dependence is deter-mined from experimental data and curve-fitting techniques. In order to understand the process of non-dimensional Pi-groups' formulation, one has to consider the dimension of every variable involved into experiment or simulation. The examples of dimensions are listed in Table.14.

Table 14. Dimensions of mechanical quantities
Quantity

| Acceleration | $L T^{-2}$ | $\mathrm{~m} \cdot \mathrm{~s}^{-2}$ |
| :--- | :--- | :--- |
| Action | $M L^{2} T^{-1}$ | $\mathrm{~kg} \cdot \mathrm{~m}^{2} \mathrm{~s}^{-1}$ |
| Angle (plane) | 1 | rad. |
| Angle (solid) | 1 | sterad. |
| Angular acceleration | $T^{-2}$ | $\mathrm{rad} \cdot \mathrm{s}^{-2}$ |
| Angular momentum | $M L^{2} T^{-1}$ | $\mathrm{~kg} \cdot \mathrm{~m}^{2} \mathrm{~s}^{-1}$ |
| Area | $L^{2}$ | $\mathrm{~m}^{2}$ |
| Curvature | $L^{-1}$ | $\mathrm{~m}^{-1}$ |
| Surface tension | $M T^{-2}$ | $\mathrm{~kg} \cdot \mathrm{~s}^{-2}$ |
| Density | $M L^{-3}$ | $\mathrm{~kg} \cdot \mathrm{~m}^{-3}$ |
| Elastic modulus | $M L^{-1} T^{-2}$ | $\mathrm{~kg} \cdot \mathrm{~m}^{-1} \mathrm{~s}^{-2}$ |
| Energy (work) | $M L^{2} T^{2}$ | J |
| Force | $M L T^{-2}$ | N |
| Frequency | $T^{-1}$ | s |
| Kinematic viscosity | $L^{2} T^{-1}$ | $\mathrm{~m}^{2} \mathrm{~s}^{-1}$ |
| Mass | $M$ | kg |
| Momentum | $M L T^{-1}$ | $\mathrm{~kg} \cdot \mathrm{~m}^{-1}$ |
| Power | $M \mathrm{~s}^{2} T^{-3}$ | W |
| Pressure | $M L^{-1} T^{-2}$ | $\mathrm{~N} \cdot \mathrm{~m}^{-2}$ |
| Time | $T$ | s |
| Velocity | $L T^{-1}$ | $\mathrm{~m} \cdot \mathrm{~s}^{-1}$ |
| Volume | $L^{3}$ | $\mathrm{~m}^{3}$ |

B. (Thermal quantities)

Enthalpy
$M L^{2} T^{2} \quad J$
Gas constant
$M L^{2} T^{2} \theta^{-1} \quad J K^{-1}$

Heat capacity per unit mass
$L^{2} T^{-1} \theta^{-1} \quad$ J.kg. ${ }^{-1} K^{-1}$

Heat capacity per unit volume
$L^{2} T^{-2} \theta^{-1} \quad$ J. $\mathrm{kg}^{-1} K^{-1}$

Internal energy
Latent heat of phase change
$M L^{-1} T^{-2} \theta^{-1} \quad$ J.m. ${ }^{-3} K^{-1}$
$M L^{2} T^{2} \quad J$
$L^{2} T^{-2} \quad J . k g^{-1}$
Quantity of heat
Temperature
Temperature gradient
$M L^{2} T^{-2}$
J

Thermal conductivity
Thermal diffusivity
$\theta$
K
$L^{-1} \theta \quad K . m^{-1}$

Heat transfer coefficient
$M T^{-3} L \theta^{-1} \quad$ W. $m^{-1} K^{-1}$
$L^{2} T^{-1} \quad m^{2} s^{-1}$
(Yarin 2012, 5)

As an example, action has the dimension $\left[M L^{2} T^{-1}\right]$, what means mass times squared length times inverse time. Note that length, time, mass, temperature, amount of substance, electric current and luminous intensity form the set of fundamental physical quantities, that allow to express all other quantities as products of powers of fundamental quantities.

To understand the application procedure of Pi -Theorem to results of simulations and experiments, one can consider next example:

Thrust of propeller " $F$ " depends on diameter " $d$ ", velocity at inlet " $v$ ", fluid density " $\rho$ ", spindle speed " $f$ " and fluid viscosity " $\mu$ ".

Involved quantities have next dimensions:

$$
\begin{gathered}
F=\left[M L T^{-2}\right] \\
d=[L] \\
v=\left[L T^{-1}\right] \\
\rho=\left[M L^{-3}\right] \\
f=\left[T T^{-1}\right] \\
\mu=\left[M L^{-1} T^{-1}\right]
\end{gathered}
$$

To determine necessary Pi-groups, one has to check the amount of fundamental quantities, used in given problem (mass, length, time $=3$ ), and then select the same amount of so-called repeating dimensionally distinct variables (they will pop up in all Pi-groups) from given list of quantities (unselected variables are called non-repeating).
There are 6 given physical quantities, knowing amount of repeating variables, one can also find the amount of Pi-groups necessary for complete description of system, using next simple formula: amount of Pigroups = amount of given physical quantities - amount of fundamental quantities involved (or Pi-groups=6-3=3 in current example).
The Pi-group is formed by multiplication of one non-repeating variable by product of unknown powers of repeating variables. (In this example one is suggested to choose following quantities as repeating variables due to their simplicity: $d, v, \rho$.) Therefore, the first Pi-group, that will be taken for " $F$ " will be:

$$
\Pi_{1}=F d^{a} v^{b} \rho^{c}
$$

In order to satisfy the dimensionality of equation (dimension of LHS = dimension of RHS) and find unknown powers "a", " $b$ " and " "", one has to replace physical quantities with their dimensions:

$$
M^{0} L^{0} T^{0}=\left[M L T^{-2}\right][L]^{a}\left[L T^{-1}\right]^{b}\left[M L^{-3}\right]^{c}
$$

Using the rule of power summation and multiplication, one will come up with system of 3 linear equations with 3 unknowns:

Mass [M]: $0=1+c$
Length [L]: $0=1+a+b-3 c$
Time [T]: $0=-2-b$
The solution such system will be: $a=-2 ; b=-2 ; c=-1$. Therefore the first Pigroup will be:

$$
\Pi_{1}=\frac{F}{\rho v^{2} d^{2}}
$$

For spindle speed " $f$ ":

$$
\begin{gathered}
\Pi_{2}=f d^{a} v^{b} \rho^{c} \\
M^{0} L^{0} T^{0}=\left[T^{-1}\right][L]^{a}\left[L T^{-1}\right]^{b}\left[M L^{-3}\right]^{c}
\end{gathered}
$$

Mass [M]: 0=c
Length [L]: $0=a+b-3 c$
Time [T]: 0=-1-b
Implying that $\mathrm{c}=0 ; \mathrm{b}=-1 ; \mathrm{a}=1$; that results to:

$$
\Pi_{2}=\frac{f d}{v}
$$

For dynamic viscosity " $\mu$ ":

$$
\begin{gathered}
\Pi_{3}=\mu d^{a} v^{b} \rho^{c} \\
M^{0} L^{0} T^{0}=\left[M L^{-1} T^{-1}\right][L]^{a}\left[L T^{-1}\right]^{b}\left[M L^{-3}\right]^{c}
\end{gathered}
$$

Mass [M]: $0=1+c$
Length [L]: $0=-1+a+b-3 c$
Time [T]: 0=-1-b
Implying that $\mathrm{c}=-1 ; \mathrm{b}=-1 ; \mathrm{a}=-1$; that results to:

$$
\Pi_{3}=\frac{\mu}{\rho v d}
$$

The group " $\Pi_{3}$ " can be replaced by its inverse without any loss of information:

$$
\frac{1}{\Pi_{3}}=\frac{\rho v d}{\mu}
$$

As one can recognize, the value " $\frac{1}{\Pi_{3}}$ " is a well-known, previously mentioned Reynolds number " $R e$ ".

Therefore, the final solution for this problem can be written as follows:

$$
\frac{F}{\rho v^{2} d^{2}}=\varphi\left(\frac{f d}{v}, R e\right)
$$

Where " $\varphi$ " unlike for chapter 2 denotes some function of two variables, that must be determined through experiments or computer simulations, and proper curve-fitting method.

Another example can be a flow passing the sphere with given variables of diameter " $d$ ", velocity at inlet " $v$ ", fluid density " $\rho$ ", drag force " $F$ " and fluid viscosity " $\mu$ ", or:

$$
\begin{gathered}
F=\left[M L T^{-2}\right] \\
d=[L] \\
v=\left[L T^{-1}\right] \\
\rho=\left[M L^{-3}\right] \\
\mu=\left[M L^{-1} T^{-1}\right]
\end{gathered}
$$

This case is similar to previous example except there is no spindle speed involved. Therefore there are only two Pi-groups necessary to derive, that will be " $\frac{F}{\rho v^{2} d^{2}}$ " and " $R e^{\prime}$. Hence the final solution will be:

$$
\frac{F}{\rho v^{2} d^{2}}=\varphi(R e)
$$

The example with propeller shows all necessary steps to form all necessary Pi-groups necessary to describe the relation between all involved variables at a certain physical process, that is the main usage of Pi-theorem in engineering, allowing to analyse massive simulation and experimental data from all physics branches (Fluid and Solid Mechanics, Electrodynamics, etc.) in reliable way.

## 5 CONCLUSION.

The material listed in this work is a sufficient minimum for students, who are just starting to explore the capabilities of computational Physics, necessary to understand, what will happen inside such solvers as FLUENT once the tutorial completing has begun and to properly justify the correct-ness of obtained results. Included FLUENT Tutorials from Cornell University on the other hand provide new students with necessary experience in application of CFD to real engineering problems. A literature for further reading after finishing of tutorials was also suggested.

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